Pseudospectra of Holographic QNMs



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D. Arean, D. Fariña, K.L. arXiv:2307.08751 [hep-th] JHEP 12 (2023) 187

Pseudospectra of Holographic QNMs

- → Quasinormal modes
- → Gauge/Gravity Duality (vulgo Holography)
- → QNMs in Holography
- → Pseudospectra
- → Pseudospectra of QNMs
- → Summary and Outlook

Normal Modes

Eigen modes of string:

$$\frac{d^2\Phi(x)}{dx^2} + \lambda\Phi(x) = 0$$

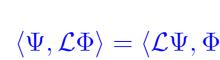
Boundary conditions:

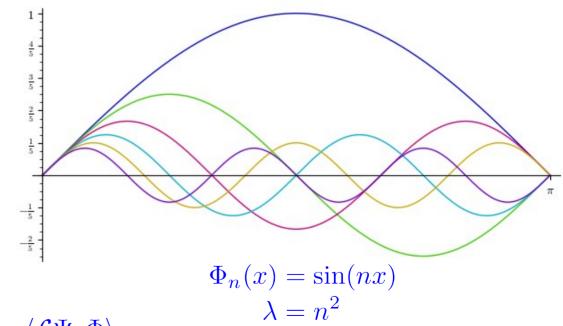
$$\Phi(0) = \Phi(\pi) = 0$$

Hermitian operator:

$$\mathcal{L} = \frac{d^2}{dx^2}$$
 , $\mathcal{L}^{\dagger} = \mathcal{L}$

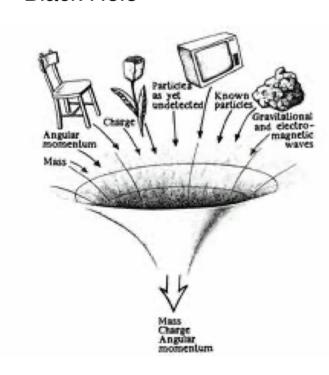
$$\langle \Psi, \Phi \rangle = \int_0^{\pi} dx \bar{\Psi} \Phi \qquad \langle \Psi, \mathcal{L}\Phi \rangle = \langle \mathcal{L}\Psi, \Phi \rangle$$





Quasi-Normal Modes

Black Hole



- → Black Holes no Hair
- → Swallow everything
- → Fate of a perturbation
 - → Either fall into BH
 - → Or radiate off to infinity
 - → Perturbation eventually dies off

Quasi-Normal Modes

How to compute Quasi Normal Modes:

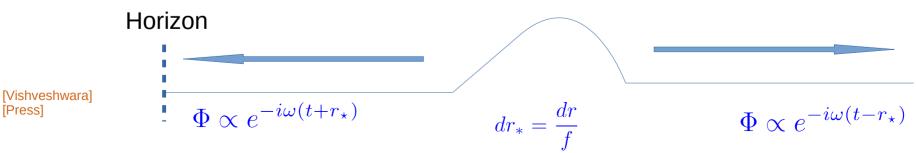
[Press]

$$ds^{2} = -f(r)dt^{2} + \frac{dr^{2}}{f(r)} + r^{2}d\Omega^{2}$$

$$\Phi = \phi(r)e^{-i\omega t}Y_{lm}(\Omega)$$

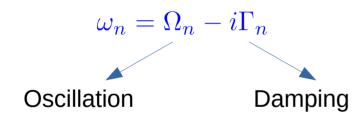
$$\phi'' + \frac{f'}{f}\phi + \left(\frac{\omega^{2}}{f^{2}} - \frac{l(l+1)}{r^{2}f} + \frac{f'}{rf}\right)\phi = 0$$

"Outgoing" boundary conditions:



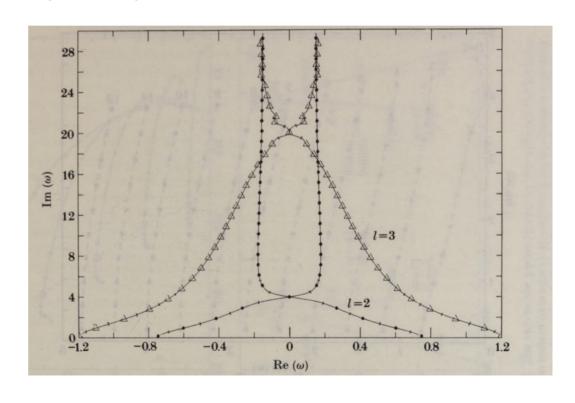
Quasi-Normal Modes

Leaky boundary conditions lead to complex frequencies

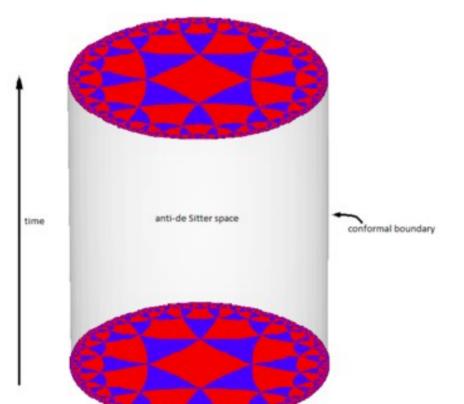


"Black Hole spectroscopy"

[Berti, Cardoso, Will, PRD 73 (2006)]



Holography



Gravity in asymptotically AdS = QFT

| Holographic Dictionary | |
|------------------------|------------------------|
| Gravity | Quantum Field Theory |
| Metric | Energy Momentum Tensor |
| Gauge field | Conserved current |
| Scalar field | Scalar operator |

Notice: Symmetry in QFT = gauge principle in AdS

Planar AdS Black hole:
$$ds^2 = \frac{r^2}{L^2} \left(-\left(1 - \pi^4 T^4/r^4\right) dt^2 + dx^2 \right) + \frac{L^2 dr^2}{r^2 \left(1 - \frac{\pi^4 T^4}{r^4}\right)} \right)$$

$$\Phi(r,t,\vec{x}) = e^{-i\omega t + i\vec{k}.\vec{x}} \phi_{\omega,\vec{k}}(r)$$

Boundary condition Horizon: $\phi_{\omega,\vec{k}} \propto e^{-i\omega(t+r_*)}$

Boundary condition boundary: $\phi \approx A(\omega, \vec{k})r^{-\Delta_{-}}(1 + \dots) + B(\omega, \vec{k})r^{-\Delta_{+}}(1 + \dots)$

Retarded Green's function:

$$G_R(\omega, \vec{k}) = K \frac{B(\omega, k)}{A(\omega, \vec{k})}$$

[Horowitz, Hubeny], [Birmingham, Sachs, Solodhukin] [Kovtun, Son, Starinets]



$$\Phi \propto r^{-\Delta_+}$$

Example: Scalar field in BTZ black hole

$$G_R(\omega, k) = \frac{(\omega^2 - k^2)}{4\pi^2} \left[\psi \left(1 - i \frac{\omega - k}{4\pi T} \right) + \psi \left(1 - i \frac{\omega + k}{4\pi T} \right) \right]$$

$$\omega_n = \pm k - i4(n+1)$$

Exact spectrum of QNMs!

In general no exact solution, e.g. scalar in AdS₅:

$$\phi'' + \left(\frac{5}{r} + \frac{f'}{f}\right)\phi' + \frac{\omega^2 - f^2\vec{k}}{r^2f}\phi = 0$$

"Christmas tree"

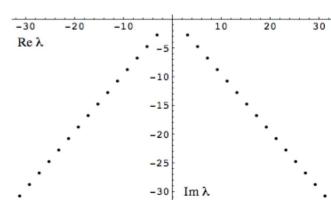


Fig. 1: The lowest 15 quasinormal frequencies in the complex λ -plane for q=0. [Starinets]

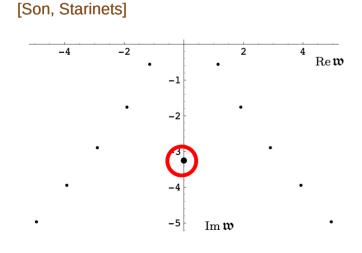
Gauge fields: new ingredient gauge symmetry: conserved current $\partial_{\mu}J^{\mu}=0$

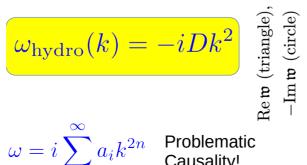
$$\frac{d}{dt}Q = 0 \qquad Q = \int d^3x J^0$$

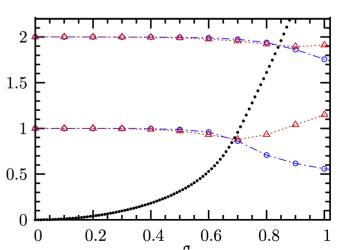
2 channels:

- Transverse is like scalar
- Longitudinal new: diffusion

[Amado, Hoyos, K.L., Montero]





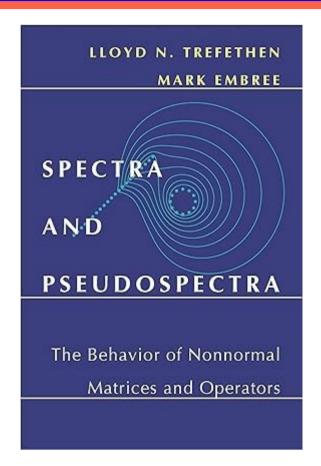


Quasinormal Modes are an essential ingredient of the Gauge/Gravity duality!

- QNMs are eigenvectors of non-Hermitian operators
- No spectral theorem $\mathcal{O} \neq \sum_{n} |n\rangle \lambda_n \langle n|$
- QNMs are not complete (only late time)
- Eigenfunctions are singular in Schwarzschild coords
- Robustness and physical significance in question

Near Horizon
$$\phi(r)=e^{-i\omega r_*}=e^{-i\Omega r_*-\Gamma r_*}$$
 , $r_*\to -\infty$

No Hilbert space interpretation



Resolvent: $\mathcal{R}(\mathcal{L}, z) = (\mathcal{L} - z)^{-1}$

Spectrum: $\sigma(\mathcal{L}) = \{z \in \mathbb{C} : "\mathcal{R}(\mathcal{L}, z) = \infty"\}$

Eigenvalues: $\mathcal{L}u_n = \lambda_n u_n$

Operator norm: $||\mathcal{L}|| = \sup_{u \in H} \frac{||\mathcal{L}u||}{||u||}$

Definitions of Pseudospectra:

1) Resolvent norm
$$\sigma_{\epsilon}(\mathcal{L}) = \{z \in \mathbb{C} : ||\mathcal{R}(\mathcal{L}, z)|| > 1/\epsilon\}$$

2) Perturbation
$$\sigma_{\epsilon}(\mathcal{L}) = \{z \in \mathbb{C}, \exists \delta \mathcal{L}, ||\delta \mathcal{L}|| < \epsilon : z \in \sigma(\mathcal{L} + \delta L)\}$$

3) Pseudo eigenvector
$$\sigma_{\epsilon}(\mathcal{L}) = \{z \in \mathbb{C}, \exists u^{\epsilon} : ||(\mathcal{L} - z)u^{\epsilon}|| < \epsilon ||u^{\epsilon}||\}$$

Theorem: The 3 definitions are equivalent

Condition number:
$$\kappa_i = \frac{||v_i|| ||u_i||}{|\langle v_i, u_i \rangle|}$$

Right eigenvector:

$$\mathcal{L}u_i = \lambda_i u_i$$

Left eigenvector: $\mathcal{L}^{\dagger}v_i = \lambda_i^* v_i$

Perturbation: $||\delta \mathcal{L}|| = \epsilon$

$$||\delta \mathcal{L}|| = \epsilon$$

Perturbed eigenvalue: $(\mathcal{L} + \delta \mathcal{L})u_i(\epsilon) = \lambda(\epsilon)u_i(\epsilon)$

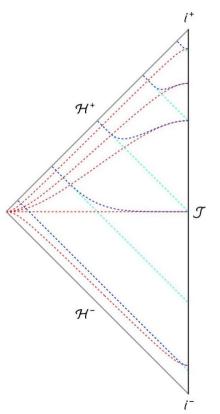
$$|\lambda(\epsilon)_i - \lambda_i| \le \epsilon \kappa_i$$

Def "small": Let d_{min} be the minimal distance between disconnected regions in the spectrum.

$$\delta \mathcal{L}$$
 , $||\delta \mathcal{L}|| = \epsilon$ is small if

$$\frac{\epsilon}{d_{min}} \ll 1$$

How to deal with the QNM problems: chose better coordinates!



Schwarzschild coordinates (worst)

$$ds^{2} = r^{2} \left[-f(r)dt^{2} + d\vec{x}^{2} \right] + \frac{dr^{2}}{r^{2}f(r)}$$
 $\phi(r) \propto (r - r_{h})^{-i\omega/2}$

Infalling Eddington-Finkelstein

$$dv = dt + \frac{dr}{r^2 f(r)}$$

$$ds^2 = r^2 \left[-f(r)dv^2 + d\vec{x}^2 \right] + 2dvdr$$

"Regular"

$$\tau = v - (1 - r_h/r)$$

$$ds^2 = r^2 \left[-f d\tau^2 + d\vec{x}^2 \right] + 2(1 - f) d\tau dr + (2 - f) \frac{dr^2}{r^2}$$

 $ightharpoonup \phi(r)$ regular at $\ r=r_h$

We need a physically motivated norm: Energy!

Energycurrent:
$$J=t^{\mu}T_{\mu\nu}dx^{\mu}$$

$$E[\Phi] = \int_{\Sigma_t} \star J$$

- We want to re-write the wave equation as a standard eigenvalue problem
- Only possible in regular coordinates

$$\psi = \partial_{\tau} \Phi$$
 $\Psi = \begin{pmatrix} \Phi \\ \psi \end{pmatrix}$ $\mathcal{L} = i \begin{pmatrix} 0 & 1 \\ L_1 & L_2 \end{pmatrix}$ \longrightarrow $\mathcal{L} \Psi = \omega \Psi$

• Compactify radial coordinate $\rho = 1 - \frac{r_h}{r}$

$$L_1 = [f(\rho) - 2]^{-1} \left[\frac{m^2 l^2}{(1 - \rho)^2} + \mathfrak{q}^2 - (1 - \rho)^3 \left(\frac{f(\rho)}{(1 - \rho)^3} \right)' \partial_\rho - f(\rho) \partial_\rho^2 \right]$$

$$L_2 = [f(\rho) - 2]^{-1} \left[(1 - \rho)^3 \left(\frac{f(\rho) - 1}{(1 - \rho)^3} \right)' + 2 (f(\rho) - 1) \partial_{\rho} \right]$$

• Adjoint operator in energy norm $\mathcal{L}^\dagger = \mathcal{L} + \begin{pmatrix} 0 & 0 \\ 0 & -i \, \delta(\rho) \end{pmatrix}$ $\frac{d}{d\tau} E = -\bar{\psi} \psi|_{\rho=0}$

- No exact solutions → numerical methods
- Pseudospectral methods
- Chebyshev polynomials for interpolation

$$F(\rho) \approx \sum_{n=0}^{N} c_n T_n(\rho)$$

$$F(\rho_j) = \sum_{n=0}^{N} c_n T_n(\rho_j) \quad , \quad \rho_j = \frac{1}{2} \left(1 - \cos \left(\frac{j\pi}{N} \right) \right) \quad , \quad j = 0 \dots N$$

- Differential operator becomes a (N+1)x(N+1) matrix D $F'(\rho_j) = D_{jk}F(\rho_j)$
- Boundary conditions: delete rows and columns corresponding to $\rho=1$ regularity corresponds to no boundary condition at $\rho=0$
- Resolvent norm becomes maximal svd $||\mathcal{L} \omega \mathbf{1}|| \approx \inf(\text{svd})$
- Energy norm becomes a metric 2Nx2N matrix $E \approx \bar{u}_k^* G_E^{km} u_m$, $u \approx (\phi(\rho_j), \psi(\rho_j))^T$

A toy example:
$$A = \begin{pmatrix} -1 & 0 \\ -50 & - \end{pmatrix}$$

 $A=\begin{pmatrix} -1 & 0 \ -50 & -2 \end{pmatrix}$ Eigenvalues: $\lambda_1=-1$, $\lambda_2=-2$

1. ℓ_2 norm: $||u|| = [\bar{u}.u]^{1/2}$

$$\kappa_1 = \kappa_2 = \sqrt{2501} \approx 50$$

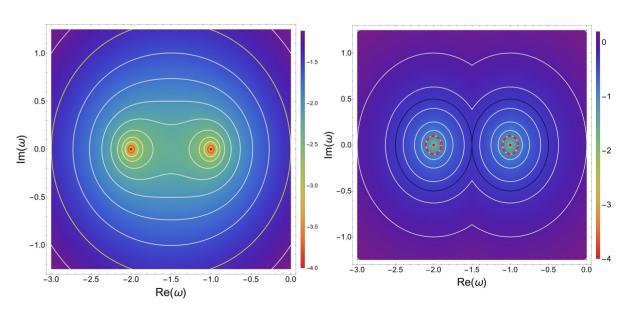
1. G-norm: $||u||_G = [\bar{u}.G.u]^{1/2}$

$$A^{\dagger} = \left[(G.A.G^{-1})^T \right]^* = A$$

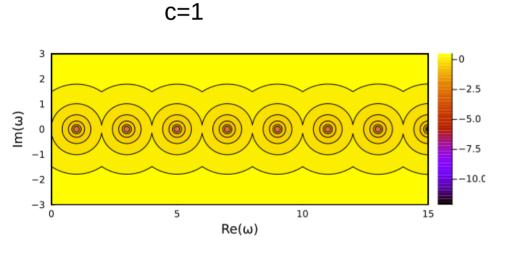
$$G = \begin{pmatrix} 20000 & 50 \\ 50 & 1 \end{pmatrix}$$

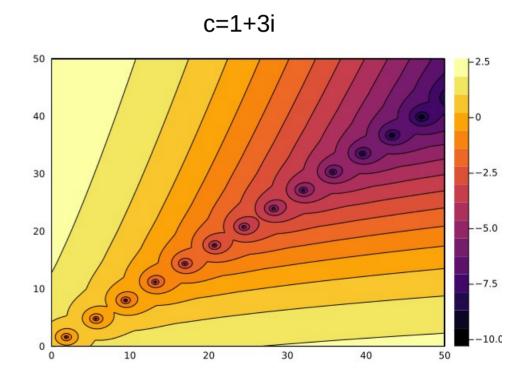
$$\kappa_1 = \kappa_2 = 1$$

Contourmaps of $\log ||A - \omega \mathbf{1}||$

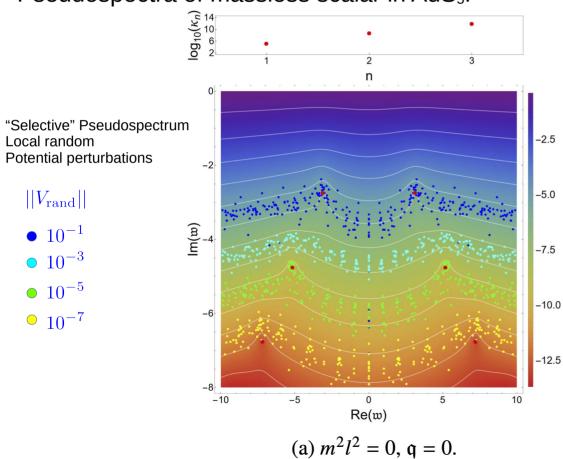


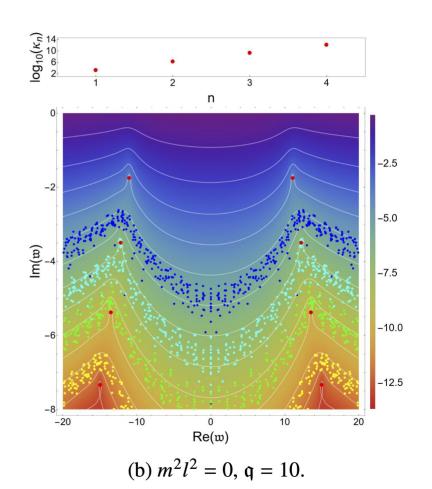
Harmonic Oscillator: $-\frac{d^2\phi}{dx^2} + c\,x^2\,\phi = \omega\phi$





Pseudospectra of massless scalar in AdS₅:



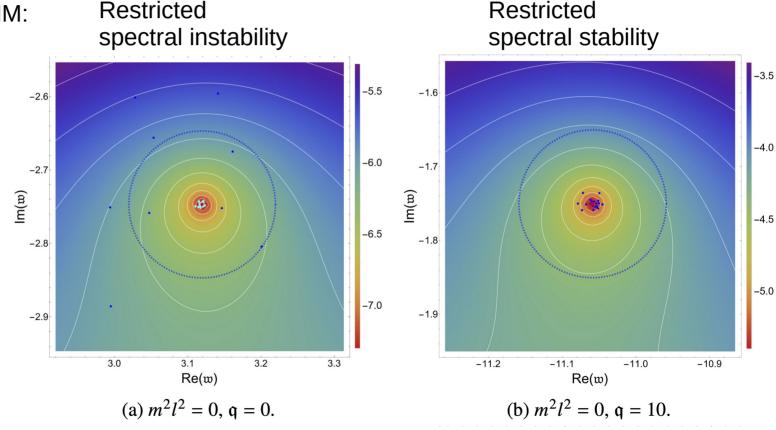


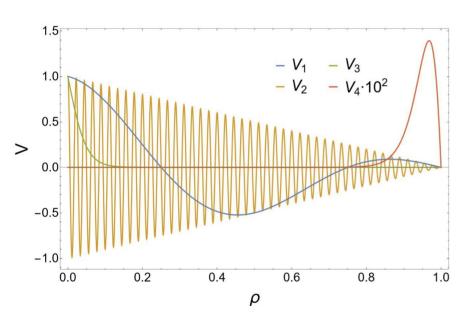
Zoom into first QNM:

Circle of stability

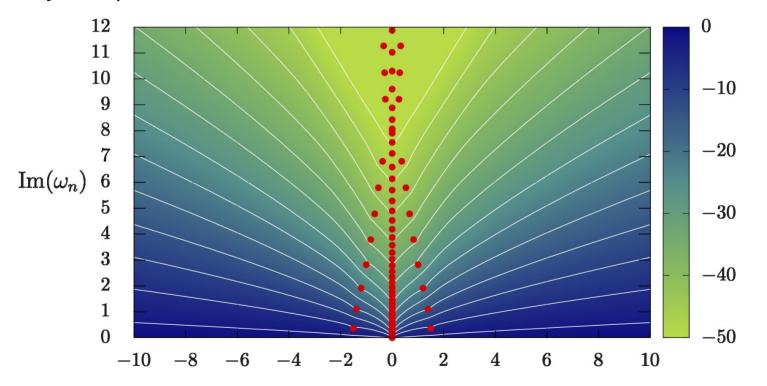
$$|\lambda(\epsilon)_i - \lambda_i| \le \epsilon \kappa_i$$

- 10^{-1}





In asymptotically flat space: [Jaramillo, Macedo, Al Sheik, PRX 11 (2021) 3, 031003 • e-Print: 2004.06434]



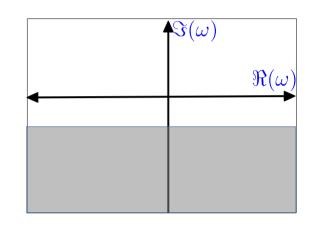
Review: [Destounis, Duque] e-Print: 2308.16227

Caveats: Ingoing modes are integrable in the energy norm:

$$\phi_{\rm in} \approx \rho^{i\omega} = \rho^{i\Omega} \rho^{\Gamma}$$

$$|\phi'_{\rm in}|^2 \approx \rho^{2\Gamma - 1} \qquad \Gamma = -\Im(\omega) > \frac{1}{2}$$

- Hilbert space of square integrable functions with energy norm In this sense the ingoing solutions belong to the spectrum
- Pseudospectrum in IEF N→ infinity limit converges to this limit
- Math: Sobolev norm, Hilbert space H^(k)
 Physics: higher derivative theories!
- In regular coords resolvent seems to be divergent in general
- Finite N provides a natural cutoff → what is the optimal N?
 e.g stretched Horizon, Planck length away from horizon
- Universality? Holographic interpretation?



$$||\Phi||^2 = \int \sum_{m=0}^k |D^m \Phi|^2$$

[Warnick: CMP. 333 (2015) 2, 959-1035 • e-Print: 1306.5760 [gr-qc]]

[Boyanov, Cardoso, Destounis, Jaramillo, Macedo] e-Print: 2312.11998 [gr-qc]

Summary

- Quasinormal Modes are central in gauge/gravity duality
- Subject to spectral instability
- Choice of norm is important → Energy norm
- Energy norm "H⁽¹⁾" Hilbert space
- Hydrodynamic modes seem safe ? $\lim_{k \to 0} \omega_{\mathrm{hydro}}(k) = 0$
- Physical significance of higher modes seems less clear
- Optimal lattice N ?
- Regular boundary conditions: beyond two derivative theories?



Thank you!