

# Solitons and non-equilibrium physics in holography

Yu Tian (田雨)

University of Chinese Academy of Sciences  
(中国科学院大学)

Jan 25, 2022

HoloTube - The Applied Holography Webinars Network

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# **Introduction**

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**Superfluid instability and dynamics**

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**Superfluid solitons and instabilities**

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**Concluding remarks**

# Discovery of solitons

- In 1843, John S. Russell observed a traveling soliton in the Union Canal in Scotland.
- Soliton solutions of the Korteweg–de Vries (KdV) equation

$$\partial_t \phi + \partial_x^3 \phi - 6 \phi \partial_x \phi = 0$$

$$\phi(x, t) = -\frac{1}{2} c \operatorname{sech}^2 \left[ \frac{\sqrt{c}}{2} (x - ct - a) \right]$$

# Solitons in Bose-Einstein condensate (BEC)

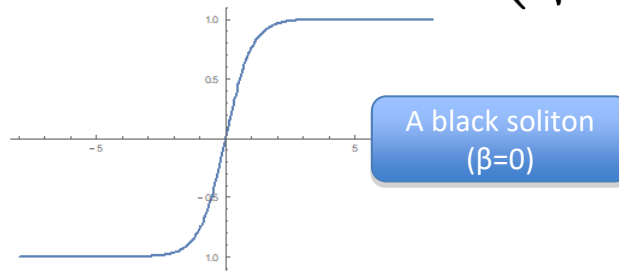
- Weakly coupled BEC described by the Gross–Pitaevskii (GP) equation ( $m = \frac{1}{2}$ , without potential)

Chemical potential

$$i\partial_t\psi = -\nabla^2\psi - (\mu - g|\psi|^2)\psi$$

- Soliton solutions of the GP equation for  $g > 0$  (repulsive): black and gray solitons

$$\psi(t, x) = \sqrt{\frac{\mu}{g}} \left[ i\beta + \sqrt{1 - \beta^2} \tanh \left( \sqrt{\frac{\mu}{2}(1 - \beta^2)}(x - \sqrt{2\mu\beta}t) \right) \right]$$

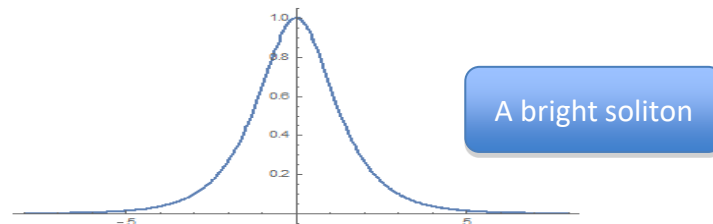


# Solitons in Bose-Einstein condensate (BEC)

- Soliton solutions of the GP equation for  $g < 0$  (attractive): bright solitons

$$\psi(t, x) = \sqrt{\frac{2\mu}{g}} \operatorname{sech}(\sqrt{-\mu}x)$$

- Bright solitons can also travel (as gray solitons for the  $g > 0$  case).



# Solitons in $\lambda\phi^4$ theory

- $\lambda\phi^4$  theory (real scalar field with double-well potential)

$$\partial_t^2 \phi - \partial_x^2 \phi = \frac{\lambda}{6} (a^2 - \phi^2) \phi$$

- Soliton solutions:

$$\phi(t, x) = a \tanh \left( \sqrt{\frac{\lambda}{12}} a \gamma (x - \beta t) \right), \quad \gamma = (1 - \beta^2)^{-1/2}$$

# Why to study solitons?

- They are simple yet nontrivial.
- They play an important role in non-equilibrium physics (e.g. in superfluids).
- They have very interesting dynamics.

# Why applied AdS/CFT (holography)?

Three key problems to be solved for (quantum) many-body systems:

- No (local) equilibrium (hydrodynamics no longer valid, e.g. near the **soliton center**)
- Strong coupling (perturbative theories no longer valid)
- Dissipation (time irreversible)

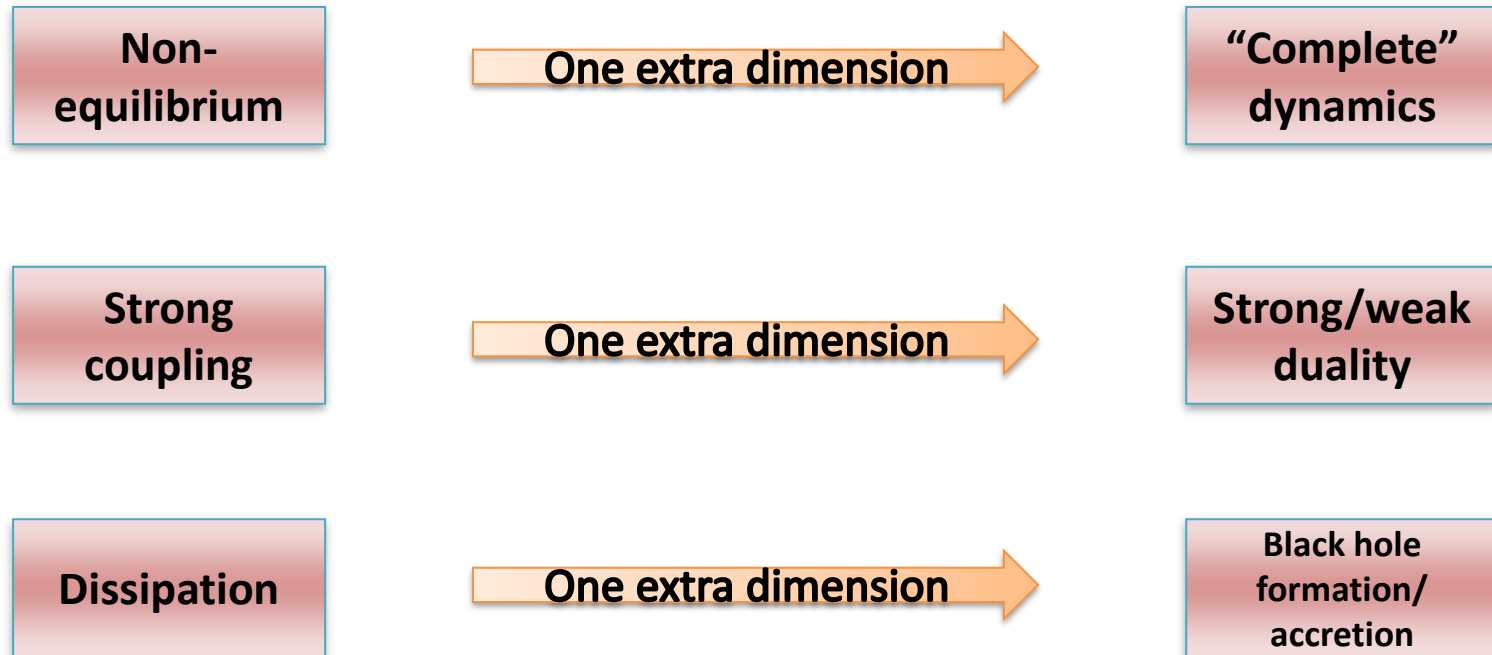
Holography deals with these problems by constructing effective field theories with **one extra dimension!**



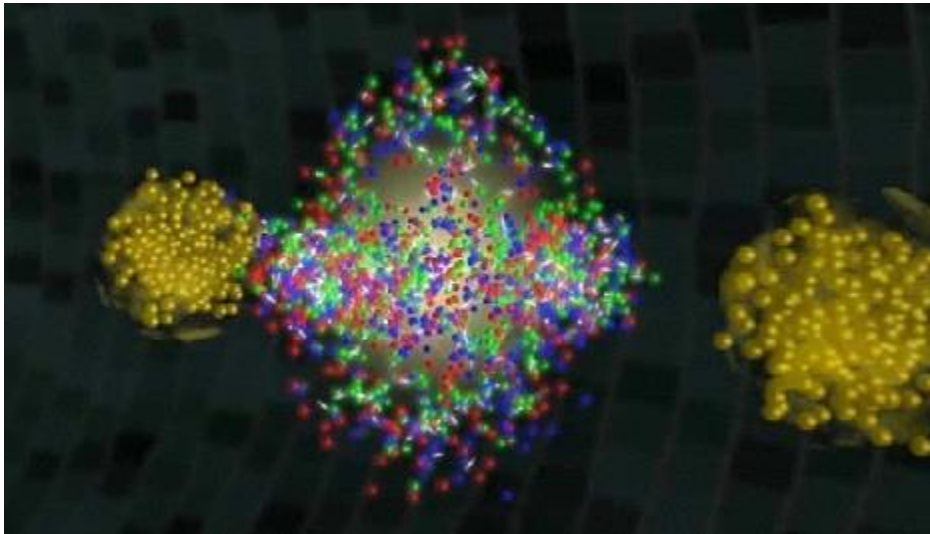
# How does holography work?

Boundary (QFT)

Bulk (AdS gravity)

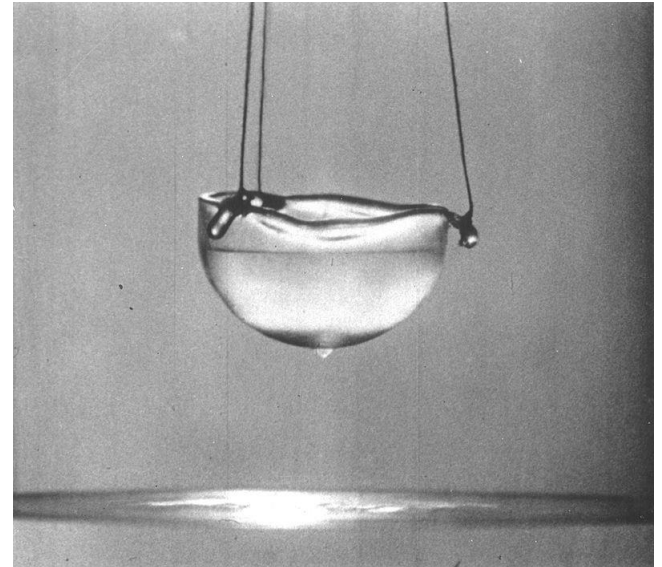


# Quantum many-body systems far from equilibrium



Quark-gluon plasma produced in LHC

The simplest holographic model: black hole formation from collision of gravitational waves in the bulk



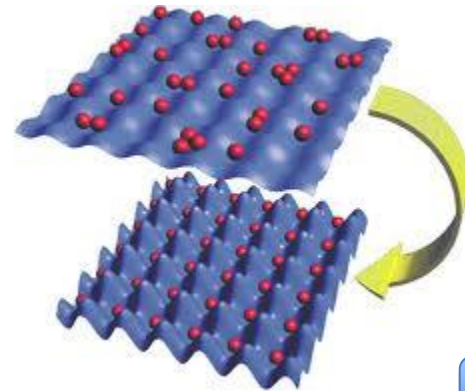
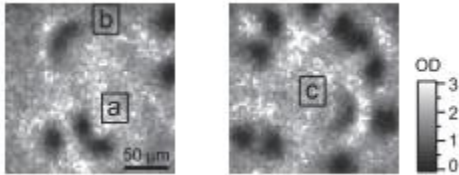
He<sub>4</sub> superfluid

U(1) symmetry breaking + gravity in the bulk (beyond Landau's two-fluid model)

# Significant progress made in experiments

Opportunities to test the holographic approach!

arXiv:  
1212.0281,  
1605.01193,  
1904.05497,  
2003.09423,  
2011.12968,  
.....



arXiv:2109.09080

Superfluid vortices in BEC

Optical lattices

(Controllable manipulation)

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## Concluding remarks

# Landau instability of the super flow

- Landau's argument (by Galilean transformations):  
When the velocity of the super flow is greater than a critical value determined by the dispersion relation of elementary excitations, the superfluid becomes unstable because excitations can have **negative** energy.
- But what is the **final** state of the superfluid if the Landau instability occurs? Breakdown of the superfluidity?

# Landau instability from holography

- In holography, elementary excitations correspond to linear perturbations on top of the bulk spacetime (and matter fields).
- A boundary system at finite temperature is dual to bulk AdS gravity with a black hole.
- Linear perturbations of the bulk black hole systems are described by the quasi-normal modes (QNM):

$$\delta\Psi(t, x, z) = e^{-i\omega t + ikx} p(z)$$

Effectively, only considering the dominant mode for stability analysis

Stable if all the quasi normal frequencies (for any  $k$ ) have negative imaginary parts

# The holographic model of a superfluid

[Hartnoll, Herzog and Horowitz, arXiv:0803.3295]

- The action:

$$D = \partial - iA, \quad q = 1$$

$$S = \int_{\mathcal{M}} d^4x \sqrt{-g} \frac{1}{q^2} \left( -\frac{1}{4} F_{ab} F^{ab} - |D\Psi|^2 - m^2 |\Psi|^2 \right)$$

- The background metric:

The most convenient choice:  $m^2 = -2/L^2$

$$ds^2 = \frac{L^2}{z^2} [-f(z) dt^2 - 2dt dz + dx^2 + dy^2], \quad f(z) = 1 - \frac{z^3}{z_h^3}$$

- The heat bath temperature:

$$T = \frac{3}{4\pi z_h}$$

- Asymptotic behavior at the AdS boundary  $z=0$

$$A_\nu = a_\nu + b_\nu z + o(z)$$

$$\Psi = \phi z + \psi z^2 + o(z^2)$$

For  $m^2 = -2/L^2$

- AdS/CFT dictionary

$\phi, \psi$ : source (turned off) & condensate (of superfluidity)

- Particle number density

$$\rho \equiv \langle J^t \rangle = -\partial_z A_t|_{z \rightarrow 0} = -b_t$$

$J^\mu$ : the only conserved current on the boundary

- Chemical potential (in equilibrium)

$$\mu = A_t|_{z \rightarrow 0} = a_t, \quad A_t(z_h) = 0$$

Radial gauge:  $A_z = 0$  all over the bulk



- Scaling symmetry

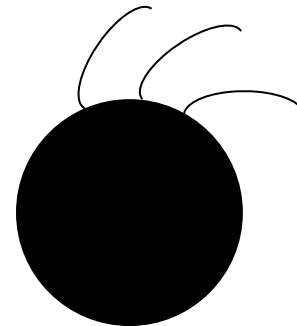
$$x^\mu \rightarrow \lambda x^\mu, \quad T \rightarrow \lambda^{-1} T, \quad \mu \rightarrow \lambda^{-1} \mu$$

Observable:  $\mu/T$   
(scaling invariant)

- The normal fluid/superfluid phase transition
  - Lowering the temperature across  $T_c$  (or increasing  $\mu$ )

Bald black hole (vanishing scalar field)	→	Hairy black hole (non-vanishing scalar)
Normal fluid (no condensate)	→	Superfluid (with condensate)
U(1) symmetry	→	SSB of U(1) (2 <sup>nd</sup> order PT)

Violation of the no-hair  
theorem in AdS)



- The super flow (along the  $x$  direction)

$$\psi \sim e^{ikx} \quad \longleftrightarrow \text{Gauge equivalent} \quad a_x = -k$$

- Velocity of the flow

$$j_\mu = \text{Im}(\psi^* \nabla_\mu \psi) \Rightarrow v^i = \frac{j^i}{j^t} = -\frac{a_i - \partial_i \theta}{\mu - \partial_t \theta} \quad (\psi = |\psi| e^{i\theta})$$

Gauge invariant!

# Landau instability from holography

- Linear analysis (QNM) of the Landau instability in holographic superfluid models  
[I. Amado et al, JHEP 1402 (2014) 063]

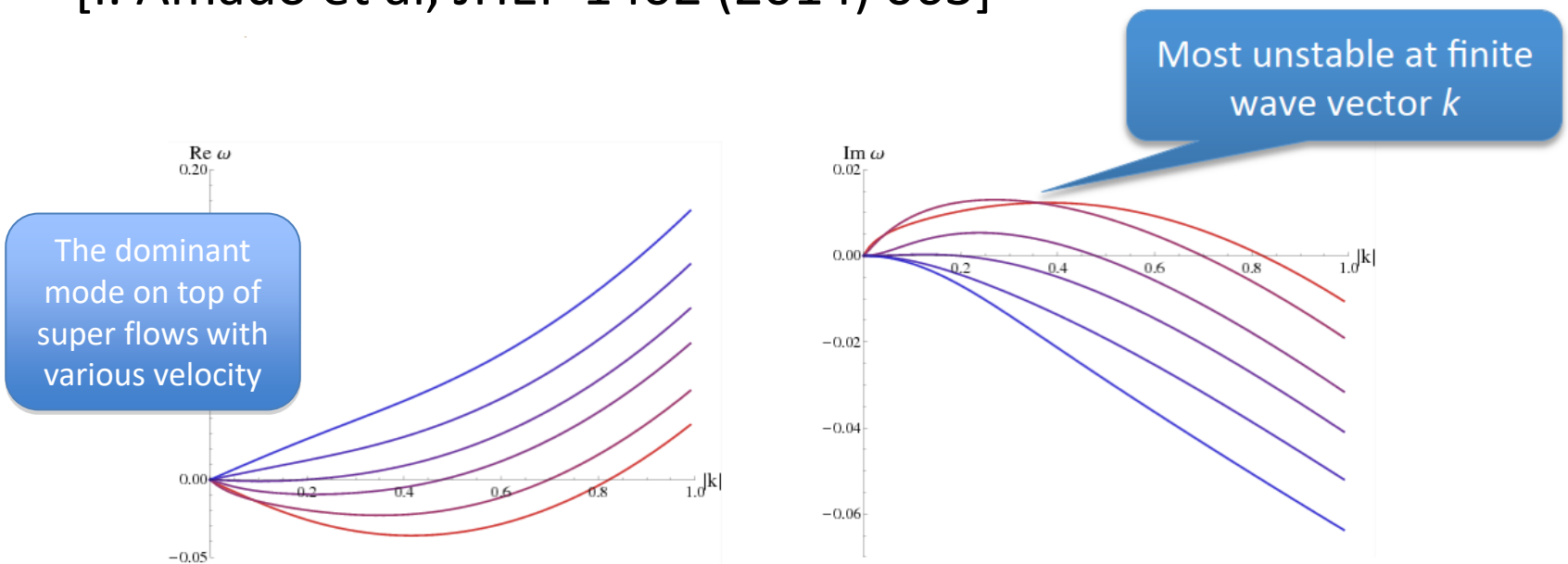


Figure: The left is for the real part of  $\omega$  and the right for the imaginary part.

- **Nonlinear dynamic evolution of the unstable super flow**
  - A too fast super flow is unstable, then what will happen? What will be **left eventually**, a striped phase (conjectured by Amado *et al*) or something else?  
[S. Lan, H. Liu, YT & H. Zhang, arXiv:2010.06232]
  - 1D: soliton generation and 1D **transient turbulence**
    - ➔ homogeneous super flow with a velocity lower than the critical value
  - 2D: soliton generation at the flow direction
    - ➔ **snake instability** of the solitons (leading to vortex nucleation) and 2D transient turbulence
    - ➔ homogeneous super flow with a velocity lower than the critical value

# Time evolution of the Landau instability in holography

The equations of motion (for numerics):

$$\partial_z^2 A_t = \partial_z \partial \cdot \mathbf{A} + i(\bar{\Phi} \partial_z \Phi - \Phi \partial_z \bar{\Phi}), \quad \Phi \equiv \Psi/z$$

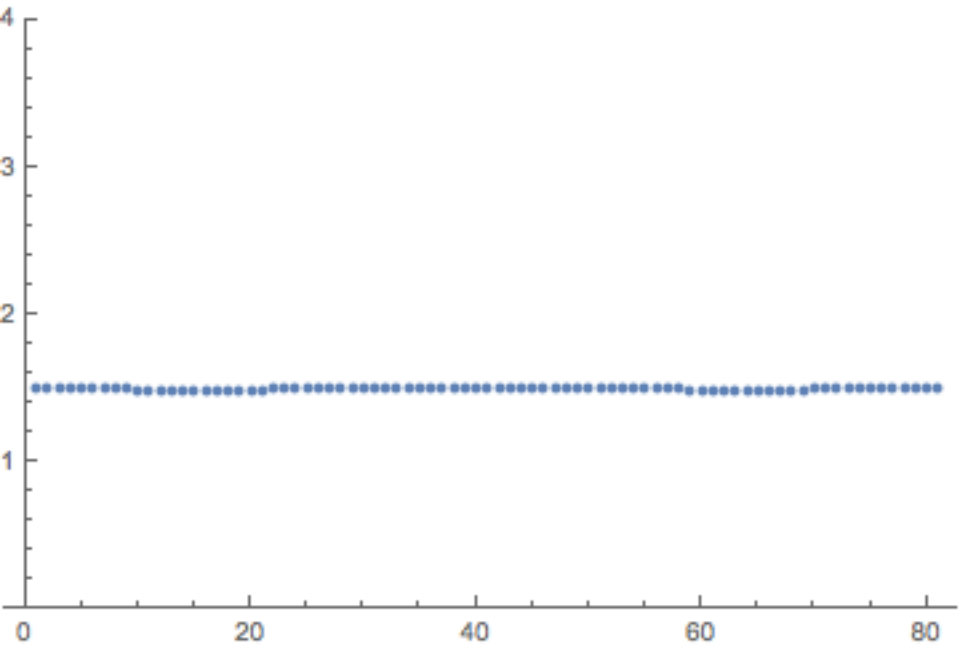
$$\begin{aligned} \partial_t \partial_z \Phi = & i A_t \partial_z \Phi + \frac{1}{2} [i \partial_z A_t \Phi + f \partial_z^2 \Phi + f' \partial_z \Phi \\ & + (\partial - i \mathbf{A})^2 \Phi - z \Phi] \end{aligned}$$

$$\begin{aligned} \partial_t \partial_z \mathbf{A} = & \frac{1}{2} [\partial_z (\partial A_t + f \partial_z \mathbf{A}) + (\partial^2 \mathbf{A} - \partial \partial \cdot \mathbf{A}) \\ & - i(\bar{\Phi} \partial \Phi - \Phi \partial \bar{\Phi})] - \mathbf{A} \bar{\Phi} \Phi \end{aligned}$$

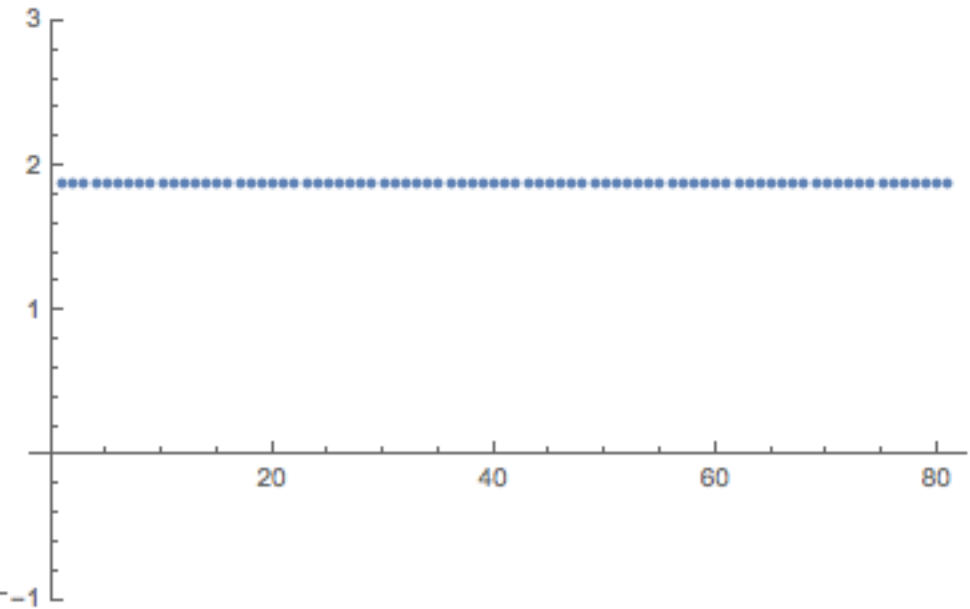
$$\begin{aligned} \partial_t \partial_z A_t = & \partial^2 A_t + f \partial_z \partial \cdot \mathbf{A} - \partial_t \partial \cdot \mathbf{A} - 2 A_t \bar{\Phi} \Phi \\ & + i f (\bar{\Phi} \partial_z \Phi - \Phi \partial_z \bar{\Phi}) - i (\bar{\Phi} \partial_t \Phi - \Phi \partial_t \bar{\Phi}) \end{aligned}$$

$$\mathbf{A} \equiv (A_x, A_y), \quad \partial \equiv (\partial_x, \partial_y)$$

- Landau instability of the super flow
  - Movies to illustrate the Landau instability in 1D



Evolution of the absolute value of the wave function (and generation of solitons)



Evolution of the super flow velocity

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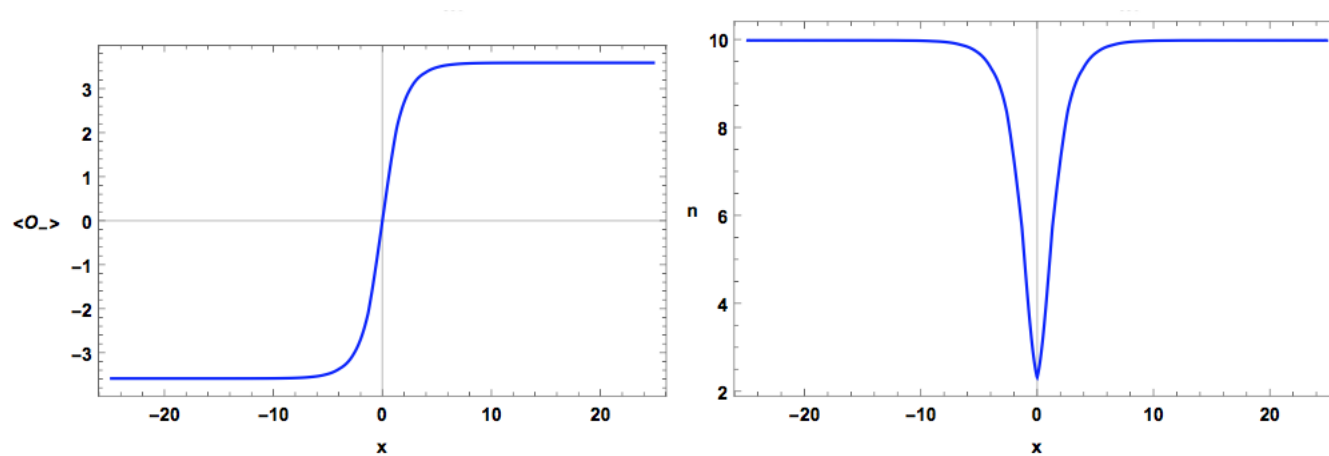
## **Superfluid solitons and instabilities**

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## **Concluding remarks**

# Black solitons in holographic superfluids at finite temperature

- Construction of holographic black solitons  
[Keranen, Keski-Vakkuri, Nowling and Yogendran, 2009; ...]

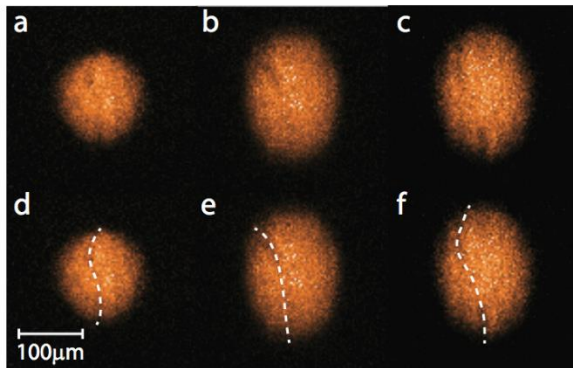


Profile of a holographic black soliton. Left: the superfluid condensate (wave function); Right: the particle number density.



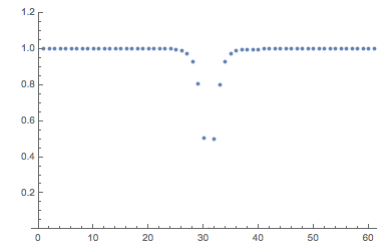
# Soliton instability and dynamic phase transition

- Snake instability: Black solitons are unstable in  $D > 1$  (has at least one transverse direction) under **bending**.



Observation of the snake instability in a fermionic superfluid [T. Yefsah *et al*, Nature 499 (2013) 426].

- Self-acceleration: Black solitons are unstable even **without bending** (or without transverse direction), in particular at finite temperature.

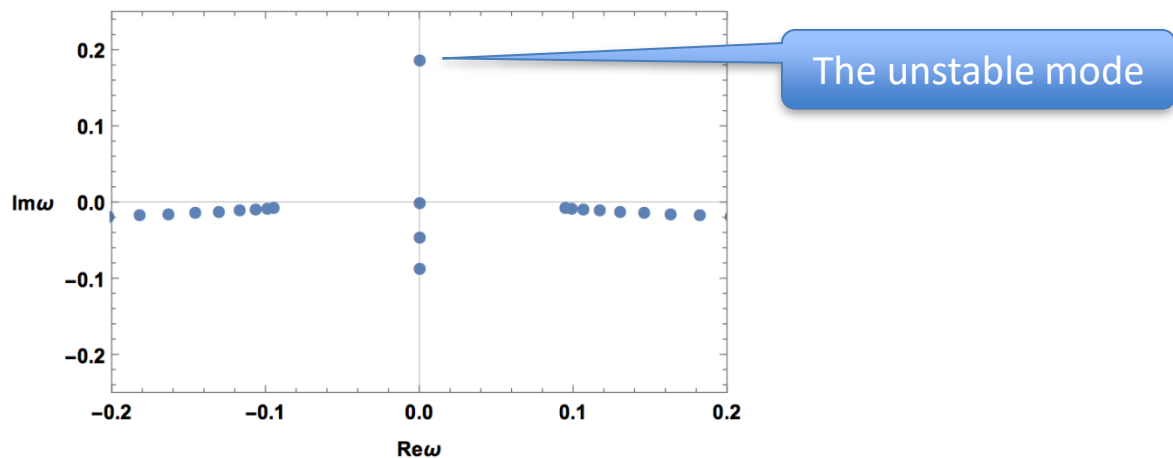


# Soliton instability and dynamic phase transition

Linear instability of black solitons from QNM of the **inhomogeneous** black hole configuration by holography  
[M. Guo, E. Keski-Vakkuri, H. Liu, YT & H. Zhang, PRL 124 (2020) 031601]

Transverse wave vector

$e^{-i\omega t + i q y}$



A typical plot of the quasi-normal modes of the holographic black soliton configuration.

# Soliton instability and dynamic phase transition

QNM on top of inhomogeneous backgrounds

- Technically complicated, so better to use the evolution-like approach (under the Eddington-Finkelstein coordinates)  
[arXiv:1511.07179](#) (homogeneous)  
[arXiv:1904.05497](#) (inhomogeneous)
- After discretization along the inhomogeneous directions, becomes a very large (generalized) eigenvalue problem

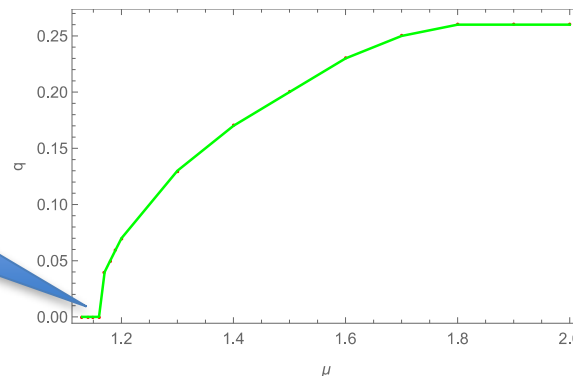
$$\delta\Psi = p(z, x)e^{-i\omega t+iqy} + \bar{p}(z, x)e^{i\omega^* t-iqy} \Rightarrow L_q(\omega) \begin{pmatrix} p \\ \bar{p}^* \end{pmatrix} = 0$$

z and x  
discretized

# Soliton instability and dynamic phase transition

- Prediction of a novel dynamical phase transition (DPT)

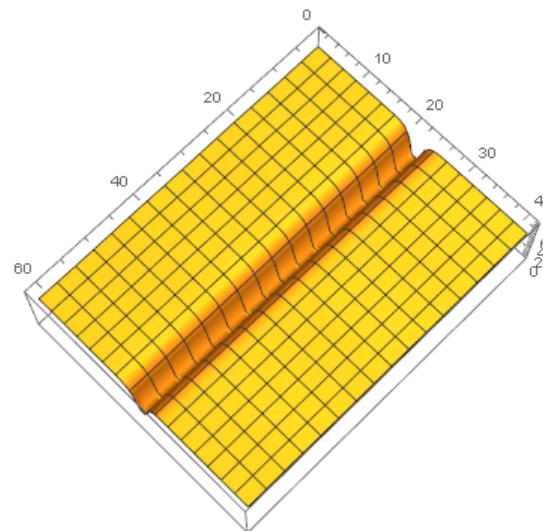
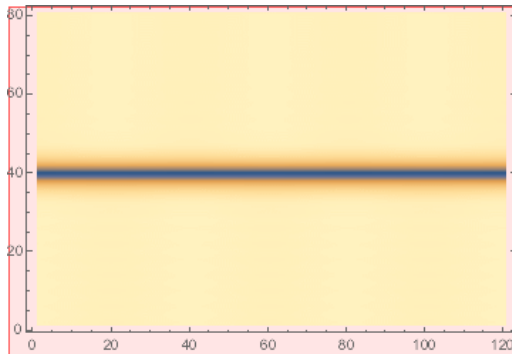
DPT at temperature  $T_d \approx 0.9T_c$  (alternative)



Plot of the transverse wave vector  $q$  of the **most unstable mode** with respect to the chemical potential  $\mu$  (or temperature  $T$ ).

$q \neq 0$ : snake instability  
 $q = 0$ : acceleration instability

- Full nonlinear evolution



# Soliton instability and dynamic phase transition

Characterization of the DPT [M. Guo, E. Keski-Vakkuri, H. Liu, YT & H. Zhang, PRL 2020]

- **Critical behavior** of the DPT as  $T \rightarrow T_d$  from below:

$$q_{\max} \propto (T_d - T)^\gamma$$

- The wave number  $q_{\max}$  determines the **vortex number density**  $n_v$  as the result of snake instability in the real time evolution:

$$n_v \propto q_{\max}$$

$n_v$  can be measured by experiments.

- In various cases,  $\gamma$  is well fit by  $\frac{1}{2}$  and a **Ginzburg-Landau-like argument** is proposed to describe the DPT and obtain  $\gamma = \frac{1}{2}$  theoretically.

# Gray solitons in holographic superfluids at zero temperature

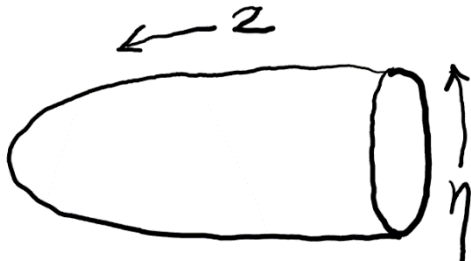
- No (holographic) gray solitons at finite temperature because of dissipation

[A. Adams, P.M. Chesler and H. Liu, Science 2012]

- Construction of holographic gray solitons at zero temperature (AdS soliton background)

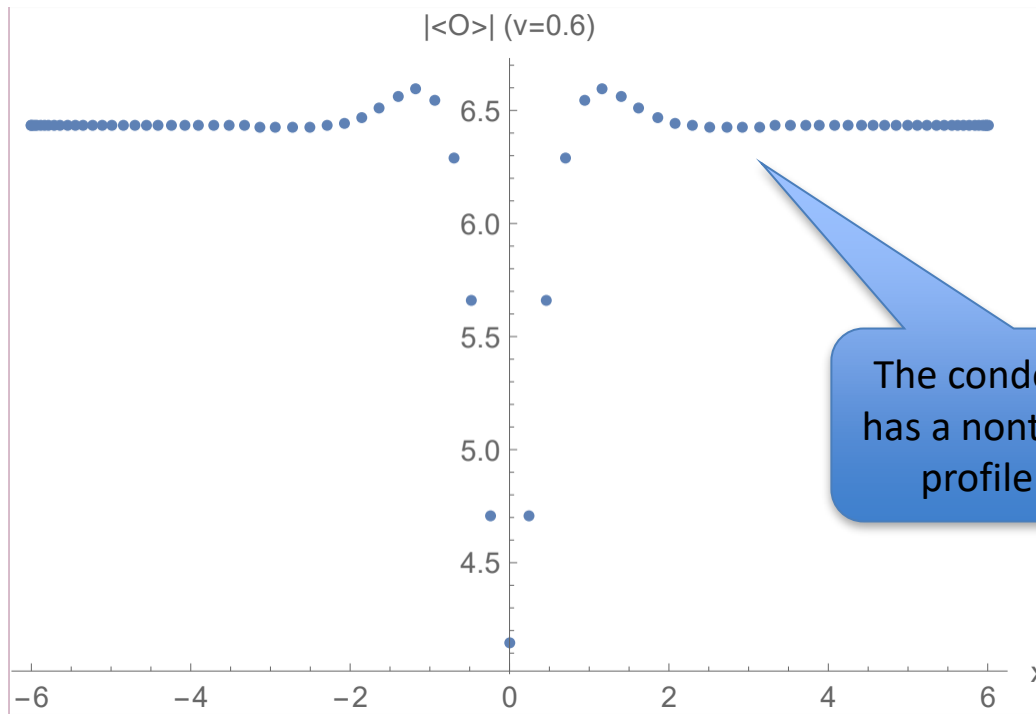
[YT *et al*, JHEP 1905 (2019) 167]

$$ds^2 = \frac{L^2}{z^2} \left[ -dt^2 + \frac{dz^2}{f(z)} + dx^2 + f(z) d\eta^2 \right], \quad f(z) = 1 - \frac{z^3}{z_0^3}.$$



Periodic to avoid the conical singularity at the tip  $z = z_0$

# Gray solitons in holographic superfluids at zero temperature



Profile of a holographic gray soliton traveling at the velocity  $v=0.6$ .

# Bright solitons in holographic superfluids





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# Conclusion

- Solitons play an important role in superfluid instability and dynamics, including quantum turbulence.
- A (superfluid) soliton is itself unstable, which leads to interesting non-equilibrium physics, like dynamic phase transition.

# Open questions

- How to characterize 1D quantum turbulence, where solitons are randomly generated?
- Bright solitons in holography?
- Gray soliton dynamics?

Thanks for your attention!

# Local equilibrium and non-equilibrium

The validity of hydrodynamics depends on **local equilibrium**, which means that the system is

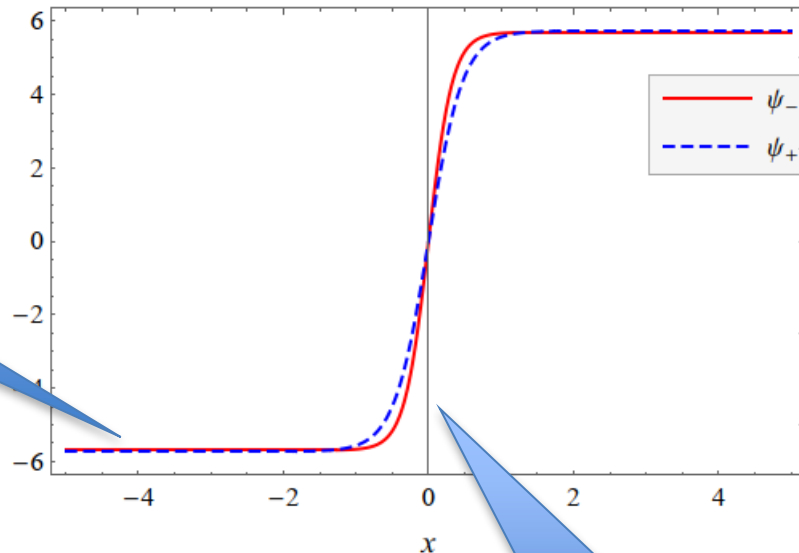
- evolving in time slowly enough
- varying in space slowly enough

with respect to all the characteristic (microscopic) time and space scales. In other words, hydrodynamics is the **low energy, long wavelength** effective theory of quantum many body systems.

For a CFT without any conserved quantities other than energy, the only characteristic scale is the temperature  $T$ , so the validity of hydrodynamics is determined by  $T$ .

# Local equilibrium and non-equilibrium

So, even a stable, **static** system can be in non-equilibrium, if there are local structures, like solitons, vortices, domain walls, etc.



Restoration of thermodynamics

Non-equilibrium region

Holographic superfluid solitons (with backreaction).  
Blue: standard quantization;  
Red: alternative quantization.  
[Z. Xu et al, arXiv:1910.09253]