Outer entropy = Bartnik-Bray quasilocal mass

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Outline

- This work brings together two notions from different contexts: Outer Entropy from holography, and Bartnik-Bray quasilocal mass from mathematical relativity, and shows their equivalence.
- Motivated by holography, it gives us some insights about an unsolved geometrical problem in geometric analysis/mathematical relativity.

The outer entropy

 The outer entropy is introduced by Engelhardt & Wall (EW) as a coarse-grained Black Hole entropy associated with an apparent horizon μ.

Engelhardt, Wall, 2018

 $\mathcal{S}(\mu) := \sup_{\rho} S_{\mathrm{vN}}(\rho) : D(\overline{\Omega}) \text{ fixed}$

- Motivated by holography, the von Neumann entropy is computed using the **Ryu-Takayanagi** surface X_{RT} , that is the extremal surface with the minimal area homologous to the boundary region.
- EW shows that for an apparent horizon μ , the maximiser always exists. $S(\mu) = \text{Area}[\mu]/4G\hbar$.
- Statistical interpretation of the BH entropy and area law. Built-in area laws associated with trapping horizons.
- Generalised by Bousso, Nomura & Remmen (BNR) to normal surfaces. They develop a EWBNR algorithm for construct fill-in data.

Nomura, Remmen, 2018 Bousso, Nomura, Remmen, 2019







Outer entropy

The outer entropy of the outer wedge data $(\overline{\Omega}, h_0, K_0)$ bounded by $\Sigma = \partial \overline{\Omega}$ with the asymptotic end B is

$$\mathcal{S}(\overline{\Omega}, h_0, K_0) := \sup_{(\Omega, h, K)} \frac{A[HRT(B)]}{4G_N \hbar} = \sup_{\substack{(\Omega, h, K) \ N \subset D(\Omega \cup_{\Sigma} \overline{\Omega}) \ \sigma \in [B]}} \max_{\substack{\sigma \in N \\ \sigma \in [B]}} \frac{\mathsf{Wall}_{[\sigma]}}{4G_N \hbar}$$

where (Ω, h, K) is the fill-in data that joins the fixed $(\overline{\Omega}, h_0, K_0)$ at Σ satisfying DEC and the following constraints:

$$\gamma|_{\Sigma_{in}} = \gamma|_{\Sigma_{out}}; \ \theta^{\pm}|_{\Sigma_{in}} = \theta^{\pm}|_{\Sigma_{out}}; \ \chi|_{\Sigma_{in}} = \chi|_{\Sigma_{out}}$$

gluing conditions

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where $\chi := K(\cdot, \ell^{-})$ is the twist or anholonomicity 1-form and ℓ^{-} is the ingoing null vector normal to Σ .



Engelhardt, Wall, 2018

Bartnik mass

- Bartnik had an idea about quasilocal mass back in the 80's. The Bartnik mass has been an active research interest in geometric analysis ever since. We seek a minimal mass extension (Ω, h, K) of initial data without horizons satisfying DEC.
- Given the positive mass theorem, the Bartnik mass $M_B(\Sigma)$ can be defined as the **infimum** ADM mass over all **horizon-free** extensions of the given surface Σ .

Bartnik, 1989

$$M_{\text{outer}}(\Sigma) := \inf_{(\overline{\Omega},h,K)} M(\overline{\Omega},h,K).$$

• Bray proposed a **dual/inner** version of the Bartnik mass in his seminal paper proving the Riemannian Penrose Inequality (RPI) :

$$M(N, h, K) \ge M_{irr}(A[\mu]) := \frac{1}{2} \left(\frac{A[\mu]}{\Omega_{n-2}} \right)^{\frac{n-3}{n-2}}.$$

$$M_{inner}(\Sigma) := \sup_{(\Omega, h, K)} \min_{\substack{\sigma \in \Omega, \\ \sigma \in [\Sigma]}} M_{irr}(A[\sigma]).$$

$$(\Omega, h, K) \xrightarrow{\Sigma} (\overline{\Omega}, h, K)$$

Equivalence

Our main result is that the outer entropy is **equivalent** to the Bartnik-Bray inner mass:

 $M_{\text{inner}}(\Sigma) = M_{\text{irr}}(4\hbar G_N \mathcal{S}(\Sigma))$

for an **outer-minimising**, mean-convex Σ .

The proof uses standard focusing arguments. Check out the details in the paper 2007.00030

A few remarks

- Σ is **outer-minimizing** means that for any Σ' enclosing Σ , $A[\Sigma] \leq A[\Sigma']$.
- Σ is **mean-convex** (normal) means that $\pm \theta^{\pm} \ge 0$.
- Both (1) outer-minimisation and (2) mean-convexity are "necessary".
- Bartnik: (1) is used to avoid "bag of gold"-like extensions trivialising the Bartnik mass.

EW: (1), as part of their "minimar" condition, is used to ensure the HRT surface can be found following their procedure.

• (2) is common in geometric analysis. e.g. Weyl problem, positivity of Brown-York mass, Liu-Yau mass, etc.

A few implications

- The area laws = monotonicity of the quasilocal mass.
- In the small sphere limit, any quasilocal mass should reduce to the stress tensor, so should the outer entropy. Calculation using the EWBNR algorithm confirms this:

$$\begin{split} \mathcal{S}(\Sigma_l) &= \frac{\Omega_{n-2} l^{n-2}}{4G_N \hbar} \left(\frac{2l^2 \Omega_{n-2} G_N T(e_0, e_0) |_p}{n-1} \right)^{\frac{n-2}{n-3}} \\ & \lim_{l \to 0} l^{-(n-1)} M_{inner}(\Sigma_l) = \frac{\Omega_{n-2} T(e_0, e_0) |_p}{n-1} \end{split}$$



Wall's ant conjecture

• A marching ant wonders what's the **minimal energy** given what she has observed so far. We look for the energy-minimising state over the purifications while **holding fixed** the interior marginal state. ρ_{Ω}

• Wall's conjecture in 1+1 dimensions:

$$\inf_{\rho: \operatorname{tr}_{\overline{\Omega}}\rho = \rho_{\Omega}} \int_{\overline{\Omega}} \langle T \rangle_{\rho} \, \mathrm{d}x = -\frac{\hbar}{2\pi} \partial_{x} S_{\rho}(\Omega) \,|_{x_{0}}, \quad \overline{\Omega} = (x_{0}, \infty), \Omega = (-\infty, x_{0}).$$
Wall, 2017

• Ceyhan-Faulkner proved it for a "null" ant on a Rindler Horizon in Minkowski spacetime w.r.t. any null variation X on a cut Σ .

$$\inf_{\rho: \operatorname{tr}_{\overline{\Omega}}\rho = \rho_{\Omega}} \int_{\Omega \cup \overline{\Omega}} \langle T \rangle_{\rho} \, \mathrm{d}x = \frac{\hbar}{2\pi} \mathscr{L}_{X} D(\rho_{\Omega} | | \sigma_{\Omega}) |_{\Sigma}.$$

Ceyhan, Faulkner, 2020

 ρ_{Ω}

•This conjecture concerns the matter sector, do we have one for the gravity sector?

Gravitational ant conjecture

• The Penrose Inequality implies $\inf_{(\overline{\Omega},h,K)} M(\overline{\Omega},h,K) \ge M_{irr}(4\hbar G_N \mathcal{S}(\Sigma)).$

• We propose the **gravitational analog** of the ant conjecture (GAC):

 $\inf_{(\overline{\Omega},h,K)} M(\overline{\Omega},h,K) = M_{\rm irr}(4\hbar G_N \mathcal{S}(\Sigma)).$

- In words, it conjectures the Bartnik mass of some closed surface is given by the irreducible mass of the largest black hole that can be fit behind it.
- It's natural that the outer entropy sits at the lower bound. It relates coarsegrained entropy with quasilocal mass, whereas Wall's ant conjecture concerns fine-grained entropy and stress tensor.
- It is a **purely geometrical** statement. Trivially, the conjecture is true if the Bartnik data can be isometrically embedded into Schwarzschild. There are also some non-trivial examples supporting this claim in the Riemannian setting.

Conclusions

- We've shown that the Bartnik-Bray quasilocal mass is equivalent to the Outer Entropy.
- We propose the gravitational ant conjecture:

$$\inf_{\overline{\Omega},h,K} M(\overline{\Omega},h,K) = M_{\rm irr}(4\hbar G_N \mathcal{S}(\Sigma)).$$

- Can we prove a Riemannian version of the conjecture ? It would be an important result in geometric analysis/mathematical relativity.
- Can we understand the microscopic origin of **any** quasilocal mass in **any** candidate theory of quantum gravity?

Thank You !

Equivalence

- We want to prove $M_{\text{inner}}(\Sigma) = M_{\text{irr}}(4\hbar G_N \mathcal{S}(\Sigma))$ for an **outer-minimising mean-convex** Σ .
- Firstly, the gluing conditions in both problems are essentially the same. They are related by a basis change in the normal bundle. In particular, $(\theta^+, \theta^-) \to (tr_{\Sigma}K, H)$.
- We need the following lemma to "quasilocalise" the outer entropy:

Lemma 1: For an outer-minimising surface Σ , the HRT surface for the outer entropy, if it exists, always lies inside the inner wedge. Nomura, Remmen, 2018

• The outer entropy is more restrictive than the inner mass:

Lemma 2: For an outer-minimising surface Σ , the supremum areas for the Bartnik-Bray inner mass and the HRT surface satisfy $A_{inner} \leq A_{HRT}$.

 ${\ensuremath{\bullet}}$ We also want $A_{inner} \leq A_{HRT} \leq A[\Sigma].$ For this, we need mean-convexity.

Lemma 3: For a mean-convex surface Σ , the HRT surface for the outer entropy, if it exists, has area $A_{HRT} \leq A[\Sigma]$. Engelhardt, Wall, 2018 Nomura, Remmen, 2018

Equivalence

- Suppose X_{HRT} , X_{inner} both exist but differ, X_{HRT} is extremal on some slice N in the interior (Lemma 1), so $A[X] \leq A[Y]$.
- The standard focusing argument, mean-convexity and DEC gives $A[X_{HRT}] \leq A[Y] \leq A[\Sigma]$, so X_{HRT} is minimal on slice $X_{HRT} Y \Sigma$.
- By definition of inner mass, $A[X_{inner}] \ge A[X_{HRT}]$. Lemma 2 also implies that $A[X_{inner}] \le A[X_{HRT}]$, so they are equal.
- Consider now the case that **neither optimiser exists**. Suppose the supremum areas satisfy $A_{HRT} > A_{inner}$, then $A_{HRT} \epsilon > A_{inner}$ for some ϵ . Area $A_{HRT} \epsilon$ is realised at some fill-in, so $A_{inner} \ge A_{HRT} \epsilon$. \Longrightarrow Contradiction.
- ${\mbox{\circ}}$ Similarly, one can show that $X_{\!H\!RT} \ \& \ X_{\!inner}$ both exist or neither exists.





Secretly an Semidefinite program?

Is the holographic dual an SDP ? SDP is a very^{$\otimes N$} important tool in QI.

 $\inf_{(\overline{\Omega},h,K)} M(\overline{\Omega},h,K) \ge M_{\operatorname{irr}}(4\hbar G_N \mathcal{S}(\Sigma))$

PI is Weak Duality: Primal \geq Dual

Primal Dual

 $\inf_{(\overline{\Omega},h,K)} M(\overline{\Omega},h,K) = M_{\operatorname{irr}}(4\hbar G_N \mathcal{S}(\Sigma))$

The gravitational ant conjecture implies the Penrose Inequality. GAC is Strong Duality: Primal = Dual