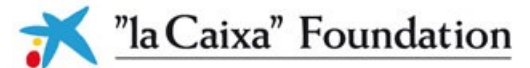


The Squashed, Stretched, Warped and Perturbed gets invaded

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HoloTube Junior
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Based on:
MC, G.Larios, O.Varela – arXiv: 2007.05172



Overview

- The Cubic deformation of ABJM and its dual GMPS
- The complete KK graviton spectrum over GMPS
- Full spectrum considerations: Space Invaders
- Conclusion and Outlook

Field Theory side: ABJM

ABJM theory is a $U(N) \times U(N)$ Superconformal Chern-Simons theory with levels $k, -k$. It is dual to a stack of M2 branes on a $\mathbb{C}^4/\mathbb{Z}_k$ orbifold. For $k = 1$ supersymmetry is enhanced to $\mathcal{N} = 8$, with $SO(8)$ bosonic subgroup.

[O.Aharony, O.Bergman,
D.Jafferis, J.Maldacena '08]

$$W = 4\pi (\mathcal{Z}^1)_{\hat{a}}^a (\mathcal{Z}^2)_{\hat{b}}^b (\mathcal{Z}^3)_{\hat{c}}^c (\mathcal{Z}^4)_{\hat{d}}^d [(\mathcal{M}^{-2})_{bc}^{\hat{a}\hat{c}} (\mathcal{M}^{-2})_{ad}^{\hat{b}\hat{d}} - (\mathcal{M}^{-2})_{bd}^{\hat{a}\hat{d}} (\mathcal{M}^{-2})_{ac}^{\hat{b}\hat{c}}] .$$

There exist two ($\mathcal{N} = 2, SU(3)$)-preserving relevant deformations of the theory:

$$\Delta W = (\mathcal{Z}^4)^p, \quad p = 2, 3,$$

leading to IR R charges

$$R_1 \equiv R(\mathcal{Z}^A) = \frac{2(p-1)}{3p}, \quad R_2 \equiv R(\mathcal{Z}^4) = \frac{2}{p} .$$

Gravity side: $\mathcal{N} = 2$ AdS geometries

$\mathcal{N} = 2$ -supersymmetric geometries are obtained as

$$d\hat{s}_{11}^2 = e^{2\Delta} \left(\frac{1}{4} ds^2(\text{AdS}_4) + ds_7^2 \right) , \quad G = \frac{m}{16} (\text{AdS}_4) + F. \quad [\text{M. Gabella, D.Martelli, A. Passias, J.Sparks '12}]$$

Two solutions of the supersymmetry constraint $\delta_\epsilon \Psi = 0$.

Necessary and sufficient conditions given in terms of spinor bilinears:

$$\begin{aligned} e^{-3\Delta} d \left[\|\xi\|^{-1} \left(\frac{m}{6} E_1 + e^{3\Delta} |S| \sqrt{1 - \|\xi\|^2} E_3 \right) \right] &= 2J_3 - 2\|\xi\| E_2 \wedge E_3 \\ d(\|\xi\|^2 e^{9\Delta} J_2 \wedge E_2) - e^{3\Delta} |S| d(\|\xi\| e^{6\Delta} |S|^{-1} J_1 \wedge E_3) &= 0 \\ d(e^{6\Delta} J_1 \wedge E_2) + e^{3\Delta} |S| d(\|\xi\| e^{3\Delta} |S|^{-1} J_2 \wedge E_3) &= 0 \end{aligned}$$

Gravity side: $\mathcal{N} = 2$ AdS geometries

The internal geometries of interest are endowed with at least two isometries, so that the line element takes the local form [M. Gabella, D.Martelli, A. Passias, J.Sparks '12]

$$ds_7^2 = \frac{f \cdot \alpha}{4\sqrt{1 + (1 + r^2)\alpha^2}} ds^2(\text{KE}_4) + \frac{\alpha^2}{16} \left[dr^2 + \frac{r^2 f^2}{1 + r^2} (d\tilde{\tau} + \sigma)^2 + \frac{1 + r^2}{1 + (1 + r^2)\alpha^2} \left(d\tilde{\psi} + \frac{f}{1 + r^2} (d\tilde{\tau} + \sigma) \right)^2 \right],$$

with f, α metric functions satisfying the ODEs

$$\frac{f'}{f} = -\frac{1}{2}r\alpha^2, \quad \frac{(r\alpha' - r^2\alpha^3)f}{\sqrt{1 + (1 + r^2)\alpha^2}} = -3.$$

For $ds^2(\text{KE}_4) = ds^2(\mathbb{CP}^2)$, the isometry group gets enlarged to $SU(3) \times U(1)_{\tilde{\psi}} \times U(1)_{\tilde{\tau}}$, with $U(1)_{\tilde{\tau}}$ broken by the four-form flux.

Gravity side: $\mathcal{N} = 2$ AdS geometries

For (at least) two choices of (f, α) the geometry can be globally extended over the S^7

$$\psi = \frac{1}{p} \tilde{\psi}, \quad \tau = \tilde{\tau} + \frac{1}{3} \left(1 - \frac{1}{p}\right) \tilde{\psi}, \quad p = 2, 3.$$

Regularity, S^7 -topology and AdS/CFT require the asymptotics

$$f \xrightarrow{r \rightarrow 0} \frac{3p}{p-1}, \quad \alpha \xrightarrow{r \rightarrow 0} w r^{-1+1/p}, \quad \text{with } w > 0,$$
$$f \xrightarrow{r \rightarrow r_0} \frac{2\sqrt{1+r_0^2}}{r_0} (r_0 - r), \quad \alpha \xrightarrow{r \rightarrow r_0} \sqrt{\frac{2}{r_0(r_0 - r)}}.$$

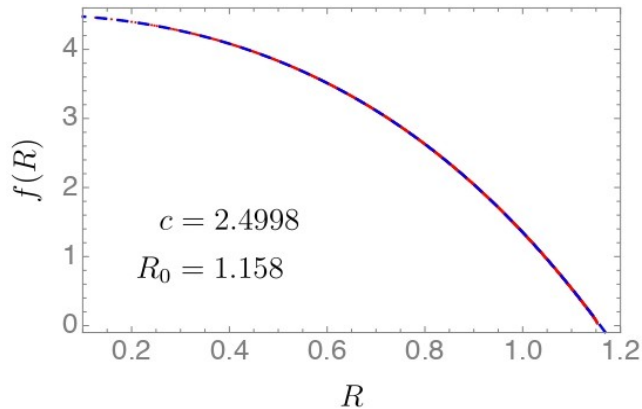
Gravity side: $p = 2$ and $p = 3$

The $p=2$ CPW solution is known analytically [R.Corrado, K.Pilch, N.Warner '01]

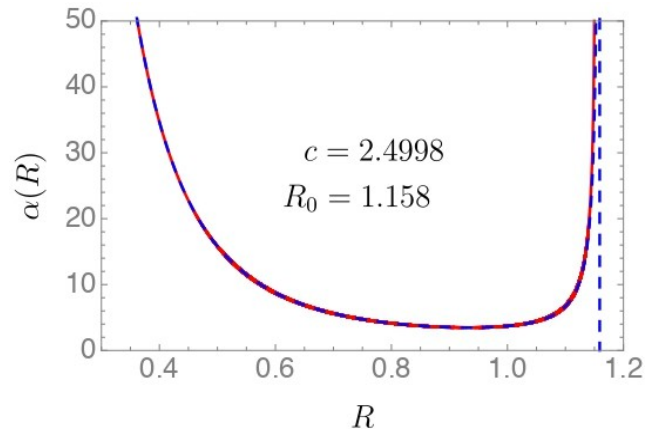
$$f = 6\left(1 - \frac{r}{r_0}\right), \quad \alpha = \sqrt{\frac{2}{r(r_0 - r)}}, \quad r_0 = 2\sqrt{2}.$$

The $p = 3$ GMPS solution is only known numerically [M. Gabella, D.Martelli, A. Passias, J.Sparks '12]

$$f(R) = \frac{9}{2} - cR^2 - \frac{c^2}{9}R^4 + \frac{(2187 - 128c^3)}{3888}R^6 + \frac{(19683c - 1264c^4)}{104976}R^8 + \mathcal{O}(R^{10}), \quad R = r^{1/3}, \quad R \in [0, R_0].$$



— Numerical Result - - - Polynomial approximation



— Numerical Result - - - Rational approximation

Kaluza-Klein Spectrum: what do we mean?

Example: One-Dimensional reduction:

$$ds^2 = ds_{\text{ext}}^2(x) + L^2 d\theta^2 .$$

Consider i.e. a massless scalar field and perform a Fourier expansion on the internal manifold

$$\phi(x, \theta) = \sum_{n=-\infty}^{\infty} \phi(x)_n \mathcal{Y}(\theta)_n = \sum_{n=-\infty}^{\infty} \phi(x)_n e^{\frac{i\pi n \theta}{L}} .$$

Then

$$\square_D \phi(x, \theta) = 0 \quad \longrightarrow \quad (\square_{D-1} + \partial_\theta^2) \phi(x, \theta) = \sum_{n=-\infty}^{\infty} \left(\square_{D-1} \phi(x)_n - \frac{\pi^2 n^2}{L^2} \phi(x)_n \right) e^{\frac{i\pi n \theta}{L}} = 0$$

Kaluza-Klein Gravitons on GMPS

Consider perturbations of the external metric

$$d\hat{s}_{11}^2 = e^{2A} \left[(\bar{g}_{\mu\nu}(x) + h_{\mu\nu}(x, y)) dx^\mu dx^\nu + d\bar{s}_7^2(y) \right], \quad e^{2A} = \frac{1}{4} e^{2\Delta}, \quad d\bar{s}_7^2 = 4 ds_7^2,$$

factorising as $h_{\mu\nu}(x, y) = h_{\mu\nu}^{[tt]}(x)\mathcal{Y}(y)$ and satisfying $\bar{\square} h_{\mu\nu}^{[tt]} = (M^2 L^2 - 2)h_{\mu\nu}^{[tt]}$.

Then 11 dimensional Einstein equations imply [C.Bachas, J.Estes '11]

$$\mathcal{L} \mathcal{Y} = L^2 M^2 \mathcal{Y}, \quad \text{with} \quad \mathcal{L} = -\frac{e^{-9A}}{\sqrt{\bar{g}_7}} \partial_m (e^{9A} \sqrt{\bar{g}_7} \bar{g}^{mn} \partial_n), \quad m, n = 1, \dots, 7.$$

Kaluza-Klein Gravitons on GMPS

$$\begin{aligned}
 \mathcal{L} = & -\frac{4}{r\alpha^2 f^3} \partial_r \left[r f^3 \partial_r \right] - \frac{\sqrt{1 + (1 + r^2)\alpha^2}}{f \cdot \alpha} \square_{S^5} && \text{[MC, G.Larios, O.Varela '20]} \\
 & -\frac{4}{9} \left(1 + \frac{1}{r^2 \alpha^2} \right) \partial_\psi^2 - \frac{8}{3} \left[\frac{2}{9} \left(1 + \frac{1}{r^2 \alpha^2} \right) - \frac{1}{r^2 \alpha^2 f} \right] \partial_\psi \partial_\tau \\
 & - \left[-\frac{\sqrt{1 + (1 + r^2)\alpha^2}}{f \cdot \alpha} + \frac{16}{81} \left(1 + \frac{1}{r^2 \alpha^2} \right) + \frac{4(1 + r^2)}{r^2 \alpha^2 f^2} - \frac{16}{9r^2 \alpha^2 f} \right] \partial_\tau^2 .
 \end{aligned}$$

Exploit the $SU(3) \times U(1) \times U(1)$ symmetry


$$\mathcal{Y} = \sum_{\ell, m, j} \xi_{\ell, m, j}(r) Y_{\ell, m}(z, \bar{z}, \tau) e^{ij\psi}$$

$$\square_{S^5} Y_{\ell, m} = -\ell(\ell + 4) Y_{\ell, m} , \quad \partial_\tau Y_{\ell, m} = im Y_{\ell, m}$$

Kaluza-Klein Gravitons on GMPS

$$\begin{aligned}
 L^2 M^2 \xi = & -\frac{4}{r\alpha^2 f^3} \frac{d}{dr} \left[r f^3 \frac{d\xi}{dr} \right] + \frac{\sqrt{1 + (1 + r^2)\alpha^2}}{f \cdot \alpha} \ell(\ell + 4) \xi \quad [\text{MC, G.Larios, O.Varela '20}] \\
 & + \frac{4}{9} \left(1 + \frac{1}{r^2 \alpha^2} \right) j^2 \xi + \frac{8}{3} \left[\frac{2}{9} \left(1 + \frac{1}{r^2 \alpha^2} \right) - \frac{1}{r^2 \alpha^2 f} \right] j m \xi \\
 & + \left[-\frac{\sqrt{1 + (1 + r^2)\alpha^2}}{f \cdot \alpha} + \frac{16}{81} \left(1 + \frac{1}{r^2 \alpha^2} \right) + \frac{4(1 + r^2)}{r^2 \alpha^2 f^2} - \frac{16}{9r^2 \alpha^2 f} \right] m^2 \xi .
 \end{aligned}$$

Exploit the $SU(3) \times U(1) \times U(1)$ symmetry

$$\mathcal{Y} = \sum_{\ell, m, j} \xi_{\ell, m, j}(r) Y_{\ell, m}(z, \bar{z}, \tau) e^{ij\psi}$$


$$\square_{S^5} Y_{\ell, m} = -\ell(\ell + 4) Y_{\ell, m} , \quad \partial_\tau Y_{\ell, m} = im Y_{\ell, m}$$

Kaluza-Klein Gravitons on GMPS

Numerics

The ODE is not analytically solvable without knowing explicitly f and α .

Strategy:

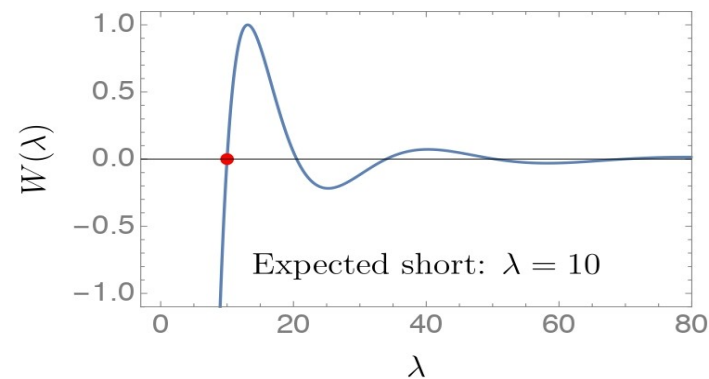
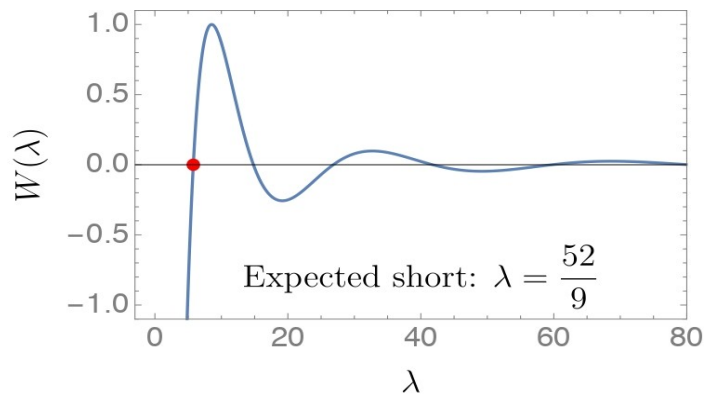
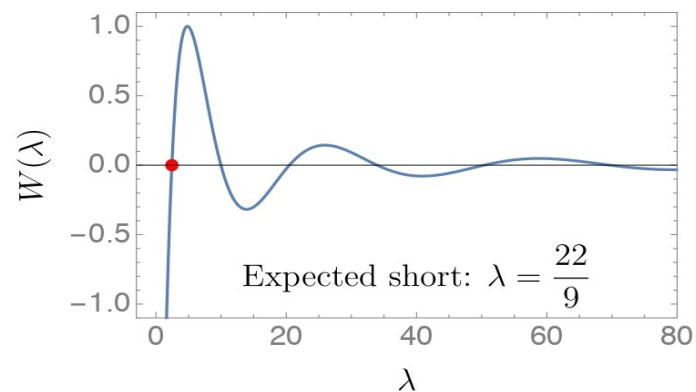
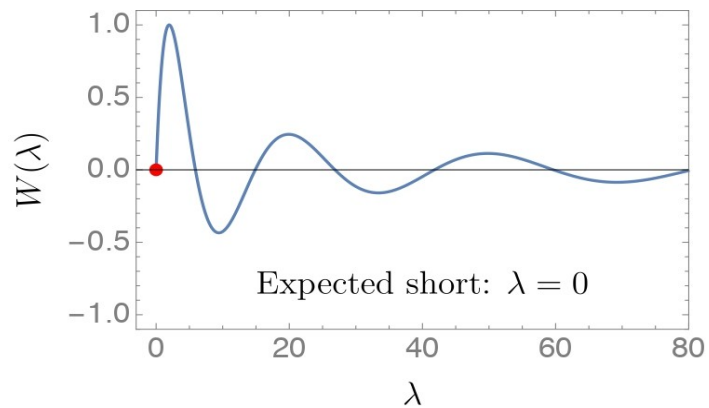
[J. Richard, R. Terrisse, D. Tsimpis '14]

- For given l, m, j find the asymptotic, normalisable values of the eigenfunctions $\xi(R)$, with $R = r^{1/3}$, close to the endpoints of $[0, R_0]$.
- Use these values as seeds to trigger a numerical integration of the ODE, starting from both endpoints of $[0, R_0]$. The resulting functions are $\xi_\lambda^L(R)$, $\xi_\lambda^R(R)$, with $\lambda = L^2 M^2$.
- The relevant eigenfunctions are the ones for which ξ_L and ξ_R are linearly dependent. This implies $W(\lambda, R) = \xi_\lambda^L(R)\dot{\xi}_\lambda^R(R) - \xi_\lambda^R(R)\dot{\xi}_\lambda^L(R) = 0$.

Kaluza-Klein Gravitons on GMPS

Numerics

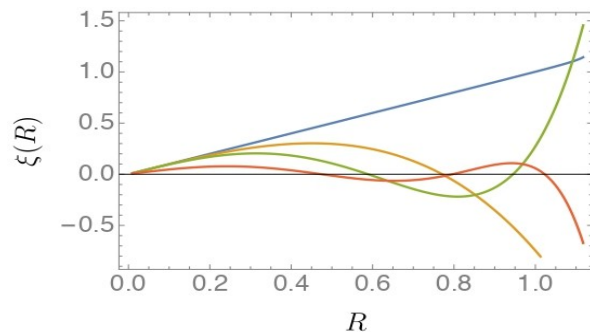
Wronskians for $l, m = 0$ and $j = 0, 1, 2, 3$.



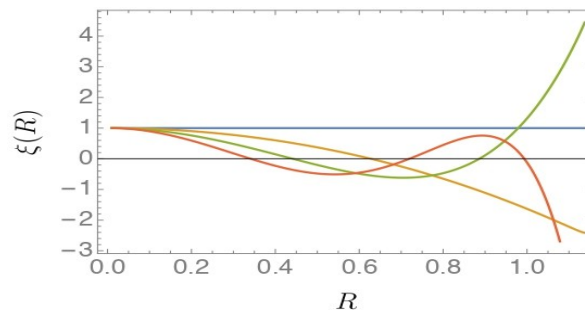
Kaluza-Klein Gravitons on GMPS

Numerics

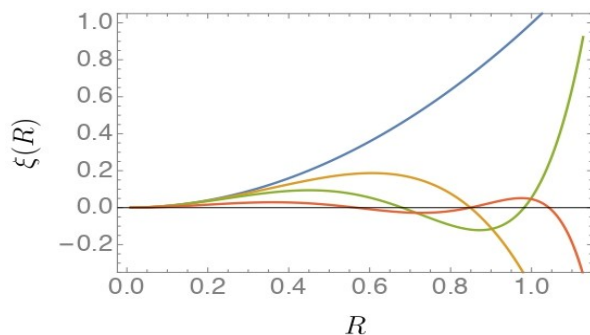
Numerical wavefunctions for $l, m = 0$ and $j = 0, 1, 2, 3$.



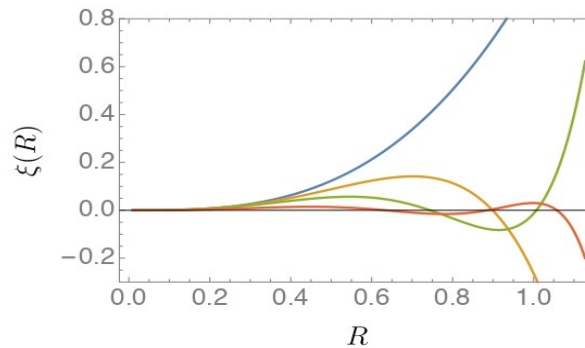
— $k=0$ — $k=1$ — $k=2$ — $k=3$



— $k=0$ — $k=1$ — $k=2$ — $k=3$



— $k=0$ — $k=1$ — $k=2$ — $k=3$



— $k=0$ — $k=1$ — $k=2$ — $k=3$

Kaluza-Klein Gravitons on GMPS

Spectrum completeness

The gravitons' $SO(8)$ representations at KK level n branch under $SU(3) \times U(1)_3$, where $U(1)_3$ such as to recover the correct $R_1 \equiv R(\mathcal{Z}^A)$, $R_2 \equiv R(\mathcal{Z}^4)$.

$$[n, 0, 0, 0] \xrightarrow{SU(3) \times U(1)_3} \bigoplus_{\ell=0}^n \bigoplus_{t=0}^{n-\ell} \bigoplus_{p=0}^{\ell} [p, \ell - p]_{-R_1(\ell-2p) + R_2(n-\ell-2t)}$$

so that, by redefining quantum numbers as

$$n = 2k + |j| + \ell, \quad m = 2p - \ell, \quad j = n - \ell - 2t$$

we are guaranteed to sweep over the complete gravitons' spectrum.

Kaluza-Klein Gravitons on GMPS

n	$[p, \ell - p]_{\frac{4}{9}(2p-\ell)+\frac{2}{3}(n-\ell-2t)}$	$d_{p, \ell-p}$	$L^2 M_{n, \ell, t, p}^2$	$\Delta_{n, \ell, t, p}$	Dual operator	Short?
0	$[0, 0]_0$	1	0	3	$\mathcal{T}_{\alpha\beta}^{(0)} _{s=2}$	✓
1	$[0, 0]_{\pm\frac{2}{3}}$	1	$\frac{22}{9}$	$\frac{11}{3}$	$\mathcal{T}_{\alpha\beta}^{(0)} \mathcal{Z}^4 _{s=2}$, c.c.	✓
	$[1, 0]_{\frac{4}{9}}, [0, 1]_{-\frac{4}{9}}$	3	1.76	3.50	$\mathcal{T}_{\alpha\beta}^{(0)} \mathcal{Z}^a _{s=2}$, c.c.	
2	$[0, 0]_{\pm\frac{4}{3}}$	1	$\frac{52}{9}$	$\frac{13}{3}$	$\mathcal{T}_{\alpha\beta}^{(0)} (\mathcal{Z}^4)^2 _{s=2}$, c.c.	✓
	$[1, 0]_{-\frac{2}{9}}, [0, 1]_{\frac{2}{9}}$	3	4.68	4.13	$\mathcal{T}_{\alpha\beta}^{(0)} \mathcal{Z}^a \bar{\mathcal{Z}}_4 _{s=2}$, c.c.	
	$[2, 0]_{\frac{8}{9}}, [0, 2]_{-\frac{8}{9}}$	6	3.88	3.97	$\mathcal{T}_{\alpha\beta}^{(0)} \mathcal{Z}^a \mathcal{Z}^b _{s=2}$, c.c.	
	$[1, 0]_{\frac{10}{9}}, [0, 1]_{-\frac{10}{9}}$	3	5.07	4.21	$\mathcal{T}_{\alpha\beta}^{(0)} \mathcal{Z}^a \mathcal{Z}^4 _{s=2}$, c.c.	
	$[0, 0]_0$	1	5.92	4.36	$\mathcal{T}_{\alpha\beta}^{(0)} (1 - 4a^2 \mathcal{Z}^4 \bar{\mathcal{Z}}_4 + b \mathcal{Z}^a \bar{\mathcal{Z}}_a) _{s=2}$	
	$[1, 1]_0$	8	4	4	$\mathcal{T}_{\alpha\beta}^{(0)} (\mathcal{Z}^a \bar{\mathcal{Z}}_b - \frac{1}{3} \delta_b^a \mathcal{Z}^c \bar{\mathcal{Z}}_c) _{s=2}$	

For $[0, 0]_{\pm R_2 n}$, the masses are well approximated by $L^2 M_n^2 = R_2 n (R_2 n + 3)$.
As $\Delta(\Delta - 3) = M^2 L^2$, their conformal dimension is locked to the R charge as

$$\Delta_n = |R_n| + 3$$

Also present in CPW [I.Klebanov, T.Klose, A.Murugan '09]

Group Theory determines the dual CFT operator:

$$\mathcal{T}_{\alpha\beta}^{(0)} (\mathcal{Z}^4)^n, \quad n = 0, 1, 2, \dots,$$

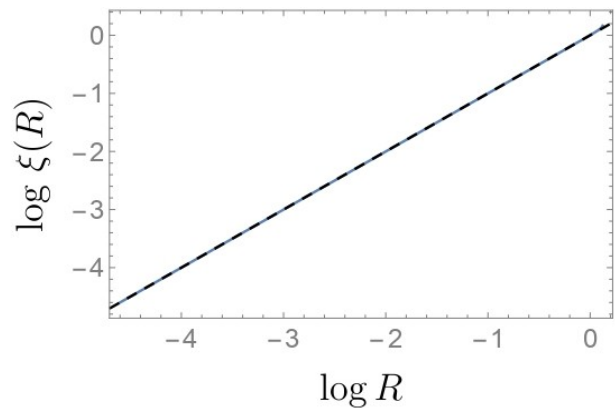
while the correspondent eigenfunction is

$$\mathcal{Y}_j = (\xi_1)^j e^{ij\psi}.$$

$$\xi_1 = r^{1/3} \equiv R$$

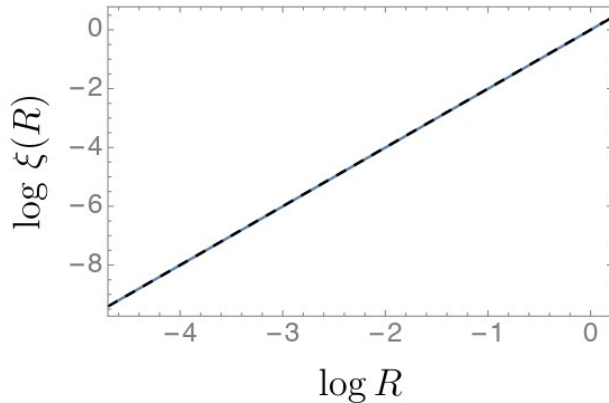
Kaluza-Klein Gravitons on GMPS

Analytics: short gravitons



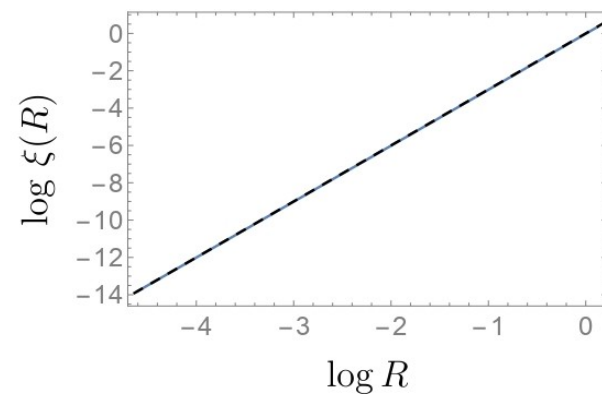
— Numerical Result -- Expected: $\xi = aR$

$$j = 1$$



— Numerical Result -- Expected: $\xi = aR^2$

$$j = 2$$



— Numerical Result -- Expected: $\xi = aR^3$

$$j = 3$$

Kaluza-Klein Gravitons on GMPS

Analytics: shadow octets

For $[1, 1]_{\pm R_2(n-2)}$, masses are well approximated by

$$L^2 M_n^2 = ((n-2)R_2 + 4)((n-2)R_2 + 1), \quad n = 2, 3, \dots \quad \Delta_n = (n-2)R_2 + 4$$

[M. Billo et al. '00]

Group Theory tells us the dual CFT operator

Also present in CPW [I.Klebanov, T.Klose, A.Murugan '09]

$$\mathcal{T}_{\alpha\beta}^{(0)} \left(Z_B^A - \frac{1}{3} \delta_B^A Z^C \bar{Z}_C \right) (Z^4)^{n-2},$$

with eigenfunction

$$\mathcal{Y}_j = \xi_8 r^{j/3} Y_{2,0} e^{ij\psi}.$$

Kaluza-Klein Gravitons on GMPS

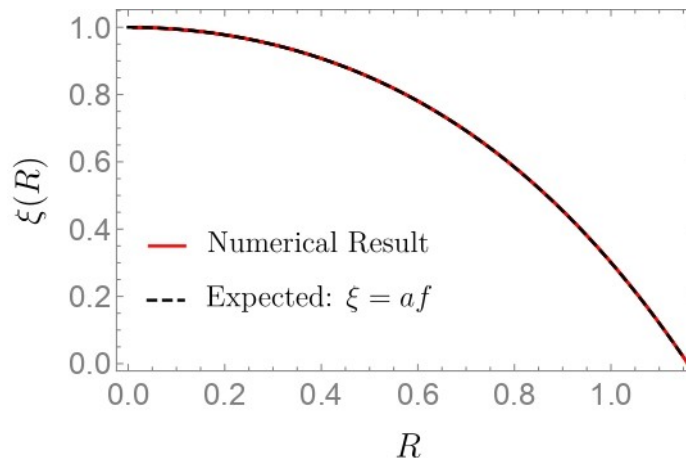
Analytics: shadow octets

The j -dependent infinite set of ODEs can be shown to amount to the set of two ODEs

$$\xi_8 + \frac{2}{r\alpha^2}\xi_8' = 0, \quad \xi_8 - \frac{3\sqrt{1+(1+r^2)\alpha^2}}{f \cdot \alpha}\xi_8 + \frac{1}{r\alpha^2 f^3}(rf^3\xi_8')' = 0.$$

! $\frac{f'}{f} = -\frac{1}{2}r\alpha^2$

$$\xi_8 \propto f$$



GMPS is not isometrically embedded in \mathbb{R}^8

If GMPS S^7 were isometrically embedded in \mathbb{R}^8 then

$$\bar{Z}_C Z^C + \bar{Z}_4 Z^4 = 1 \quad (\xi_3)^2 + (\xi_1)^2 = 1$$

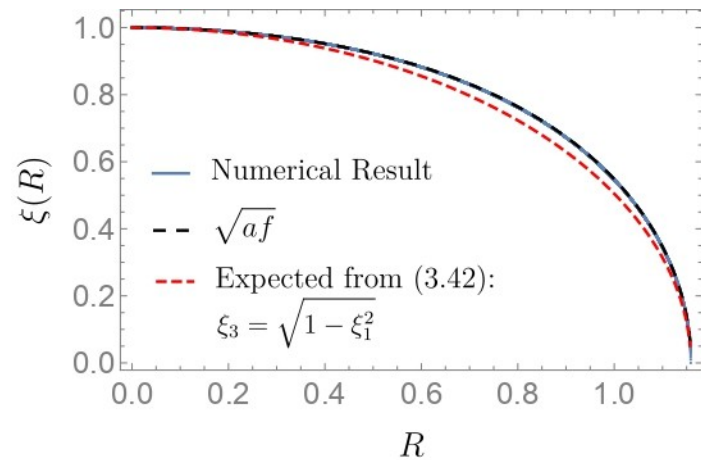
$$\xi_8 \propto f = \frac{9}{2} \left(1 - \left(\frac{r}{r_0} \right)^{2/3} \right) \quad \longrightarrow \quad \alpha^2 = \frac{4}{3r^2 \left[\left(\frac{r}{r_0} \right)^{-2/3} - 1 \right]}$$

Although satisfying the required asymptotics, there is no real value of r_0 for which GMPS ODEs are solved.

Remark:

This is not the only example of compactification on such a kind of manifold.

[M.Awada, M.Duff, C.Pope '83]



Full Spectrum considerations: Space Invaders

One could branch all the $SO(8)$ representations under $SU(3) \times U(1)_3$ and try to allocate in $OSp(4|2)_3$ supermultiplets states with different spins and same $SU(3)$ representation. For CPW [I.Klebanov, T.Klose, A.Murugan '09]

$$\text{gravitini} : \mathbf{8}_s \xrightarrow{SU(3) \times U(1)_R} \mathbf{3}_{\frac{1}{9}} + \mathbf{1}_1 + \text{''c.c.}} ,$$

$$\text{vectors} : \mathbf{28} \xrightarrow{SU(3) \times U(1)_R} \mathbf{8}_0 + \mathbf{3}_{-\frac{2}{9}} + \mathbf{3}_{-\frac{8}{9}} + \mathbf{3}_{\frac{10}{9}} + 2 \cdot \mathbf{1}_0 + \text{''c.c.}} ,$$

$$\text{fermions} : \mathbf{56}_s \xrightarrow{SU(3) \times U(1)_R} \mathbf{8}_1 + \mathbf{6}_{-\frac{1}{9}} + 2 \cdot \mathbf{3}_{\frac{1}{9}} + \mathbf{3}_{-\frac{11}{9}} + \mathbf{3}_{\frac{7}{9}} + \mathbf{1}_1 + \mathbf{1}_{\frac{1}{3}} + \text{''c.c.}} ,$$

$$\text{scalars} : \mathbf{35}_v \xrightarrow{SU(3) \times U(1)_R} \mathbf{8}_0 + \mathbf{6}_{\frac{8}{9}} + \mathbf{3}_{\frac{10}{9}} + \mathbf{3}_{-\frac{2}{9}} + \mathbf{1}_0 + \mathbf{1}_{\frac{4}{3}} + \text{''c.c.}} ,$$

$$\text{pseudoscalars} : \mathbf{35}_c \xrightarrow{SU(3) \times U(1)_R} \mathbf{8}_0 + \mathbf{6}_{-\frac{10}{9}} + \mathbf{3}_{-\frac{8}{9}} + \mathbf{3}_{-\frac{2}{9}} + \mathbf{1}_0 + \mathbf{1}_{\frac{2}{3}} + \text{''c.c.}} .$$

The Space Invaders Scenario

Spin	SO(8)	SU(3) × U(1) ₃												
2	1	1 ₀												
$\frac{3}{2}$	8_s	1 ₊₁ 1 ₋₁	3 _{$\frac{1}{9}$} 3 _{$-\frac{1}{9}$}	3 _{$-\frac{1}{9}$} 3 _{$-\frac{10}{9}$}										
1	28	1 ₀	3 _{$-\frac{8}{9}$} 3 _{$\frac{10}{9}$}	3 _{$\frac{8}{9}$} 3 _{$-\frac{10}{9}$}	8 ₀	3 _{$-\frac{2}{9}$} 3 _{$\frac{2}{9}$}	3 _{$\frac{2}{9}$} 3 _{$\frac{2}{9}$}	1 ₀						
			3 _{$-\frac{8}{9}$} 3 _{$\frac{10}{9}$}	3 _{$\frac{8}{9}$} 3 _{$-\frac{10}{9}$}										
$\frac{1}{2}$	56_s		3 _{$\frac{1}{9}$} 3 _{$-\frac{1}{9}$}	3 _{$-\frac{1}{9}$} 3 _{$-\frac{1}{9}$}	8 ₊₁ 8 ₋₁	3 _{$\frac{7}{9}$} 3 _{$-\frac{11}{9}$}	3 _{$-\frac{7}{9}$} 3 _{$\frac{11}{9}$}	1 ₊₁ 1 ₋₁	6 _{$-\frac{1}{9}$} 6 _{$\frac{1}{9}$}	6 _{$\frac{1}{9}$} 6 _{$-\frac{1}{9}$}	1 _{$\frac{1}{3}$} 1 _{$-\frac{1}{3}$}	1 _{$-\frac{1}{3}$} 1 _{$\frac{1}{3}$}		
			3 _{$-\frac{17}{9}$} 3 _{$\frac{17}{9}$}	3 _{$\frac{17}{9}$} 3 _{$-\frac{17}{9}$}		3 _{$\frac{7}{9}$} 3 _{$-\frac{11}{9}$}	3 _{$-\frac{7}{9}$} 3 _{$\frac{11}{9}$}	1 ₊₁ 1 ₋₁					3 _{$-\frac{1}{9}$} 3 _{$-\frac{1}{9}$}	
0	35_v				8 ₀	3 _{$-\frac{2}{9}$} 3 _{$\frac{2}{9}$}	3 _{$\frac{2}{9}$} 3 _{$\frac{2}{9}$}	1 ₀	6 _{$\frac{8}{9}$} 6 _{$-\frac{8}{9}$}	6 _{$-\frac{8}{9}$} 6 _{$\frac{8}{9}$}	1 _{$\frac{4}{3}$} 1 _{$-\frac{4}{3}$}	1 _{$-\frac{4}{3}$} 1 _{$\frac{4}{3}$}	3 _{$\frac{10}{9}$} 3 _{$-\frac{10}{9}$}	
	35_c		3 _{$-\frac{8}{9}$} 3 _{$\frac{8}{9}$}	3 _{$\frac{8}{9}$} 3 _{$-\frac{8}{9}$}	8 ₀	3 _{$-\frac{2}{9}$} 3 _{$\frac{2}{9}$}	3 _{$\frac{2}{9}$} 3 _{$\frac{2}{9}$}	1 ₀	6 _{$-\frac{10}{9}$} 6 _{$\frac{10}{9}$}	6 _{$\frac{10}{9}$} 6 _{$-\frac{10}{9}$}	1 _{$-\frac{2}{3}$} 1 _{$\frac{2}{3}$}	1 _{$\frac{2}{3}$} 1 _{$-\frac{2}{3}$}		
						3 _{$\frac{16}{9}$} 3 _{$-\frac{16}{9}$}	3 _{$-\frac{16}{9}$} 3 _{$\frac{16}{9}$}	1 ₀ 1 ₊₂ 1 ₋₂					3 _{$-\frac{8}{9}$} , 3 _{$\frac{8}{9}$} 3 _{$-\frac{8}{9}$} , 3 _{$\frac{8}{9}$} 3 _{$-\frac{16}{9}$} , 3 _{$\frac{16}{9}$} 1 ₀	
		Massless graviton	Short gravitino	Short gravitino	Massless vector	Short vector	Short vector	Long vector	Massive hyper	Massive hyper	Massive hyper	Massive hyper	Eaten modes	

Summary and Outlook

Summary:

- We computed numerically the full KK graviton spectrum over GMPS.
- We provided analytic results about short states and shadow octets.
- We showed that GMPS S^7 is not isometrically embedded in \mathbf{R}^8 .
- Furthermore, the full KK spectrum over GMPS displays *space invasion*.

Yet to be addressed:

- Look for an isometric embedding for the GMPS S^7 .
- Correct invasion pattern.
- No consistent uplift from D=4 $\longleftrightarrow \nexists f : S^7 \hookrightarrow R^8 \longleftrightarrow$ Space Invaders.

$SO(8)$ representations of fields at Kaluza-Klein level n

graviton : $G_n \equiv [n, 0, 0, 0]$,

gravitini : $\mathcal{G}_n \equiv [n, 0, 0, 1] \oplus [n - 1, 0, 1, 0]$,

vectors : $V_n \equiv [n, 1, 0, 0] \oplus [n - 1, 0, 1, 1] \oplus [n - 2, 1, 0, 0]$,

fermions : $\mathcal{F}_n \equiv [n + 1, 0, 1, 0] \oplus [n - 1, 1, 1, 0] \oplus [n - 2, 1, 0, 1] \oplus [n - 2, 0, 0, 1]$,

scalars : $S_n^+ \equiv [n + 2, 0, 0, 0] \oplus [n - 2, 2, 0, 0] \oplus [n - 2, 0, 0, 0]$,

pseudoscalars : $S_n^- \equiv [n, 0, 2, 0] \oplus [n - 2, 0, 0, 2]$,

Lagrangian of the cubic deformation

The effect of the cubic deformation of the superpotential at the Lagrangian level is to introduce interaction terms:

$$\mathcal{L} = \frac{1}{2} |\alpha|^2 (Z^4)^2 (\bar{Z}_4)^2 + \frac{1}{2} \alpha \chi^4 \chi^4 Z^4 + h.c. ,$$

with relevant dimensions $\Delta = 2$ and $\Delta = \frac{5}{2}$ and whose $SO(8)$ representations are the $\mathbf{294}_v$ and the $\mathbf{224}_{cv}$.