

Braneworld gravity = infinite derivative gravity

Pablo A. Cano

Universitat de Barcelona

based on arXiv: 2310.09333 w/ Aguilar, Bueno, Hennigar, Llorens
+ upcoming work w/ Bueno and Hennigar

Ávila, 16-17 November 2023



- 1 HIGHER-DERIVATIVE GRAVITY WITH COVARIANT DERIVATIVES
- 2 INDUCED GRAVITY ON THE BRANE
- 3 NONLOCAL MASSIVE GRAVITY IN $D = 3$
- 4 CONCLUSIONS

$$I = \frac{1}{16\pi G} \int d^D x \sqrt{|g|} \mathcal{L}(g^{ab}, R_{abcd}, \nabla_a)$$

Linearized equations

Constant curvature background of curvature $\Lambda = -\chi/\ell^2$

$$g_{ab} = \bar{g}_{ab} + h_{ab}$$

Linearized equations

Constant curvature background of curvature $\Lambda = -\chi/\ell^2$

$$g_{ab} = \bar{g}_{ab} + h_{ab}$$

Only linear symmetric tensors

$$\square^n R_{ab}, \quad g_{ab} \square^n R, \quad \square^n \nabla_a \nabla_b R$$

Linearized equations

Constant curvature background of curvature $\Lambda = -\chi/\ell^2$

$$g_{ab} = \bar{g}_{ab} + h_{ab}$$

Only linear symmetric tensors

$$\square^n R_{ab}, \quad g_{ab} \square^n R, \quad \square^n \nabla_a \nabla_b R$$

General form of linearized EOMs

$$\mathcal{E}_{ab} \equiv \sum_{l=0} \ell^{2l} \left[\alpha_l \square^l G_{ab}^{(1)} + \beta_l \square^l R^{(1)} \bar{g}_{ab} + \gamma_{l+1} \ell^2 \square^l [\bar{g}_{ab} \square - \bar{\nabla}_a \bar{\nabla}_b] R^{(1)} \right] = 0$$

Linearized equations

Constant curvature background of curvature $\Lambda = -\chi/\ell^2$

$$g_{ab} = \bar{g}_{ab} + h_{ab}$$

Only linear symmetric tensors

$$\square^n R_{ab}, \quad g_{ab} \square^n R, \quad \square^n \nabla_a \nabla_b R$$

General form of linearized EOMs

$$\mathcal{E}_{ab} \equiv \left[f_1(\ell^2 \bar{\square}) G_{ab}^{(1)} + f_2(\ell^2 \bar{\square}) R^{(1)} \bar{g}_{ab} + f_3(\ell^2 \bar{\square}) [\bar{g}_{ab} \bar{\square} - \bar{\nabla}_a \bar{\nabla}_b] R^{(1)} \right] = 0$$

Linearized equations

Constant curvature background of curvature $\Lambda = -\chi/\ell^2$

$$g_{ab} = \bar{g}_{ab} + h_{ab}$$

Only linear symmetric tensors

$$\square^n R_{ab}, \quad g_{ab} \square^n R, \quad \square^n \nabla_a \nabla_b R$$

General form of linearized EOMs

$$\mathcal{E}_{ab} \equiv \left[f_1(\ell^2 \bar{\square}) G_{ab}^{(1)} + f_2(\ell^2 \bar{\square}) R^{(1)} \bar{g}_{ab} + f_3(\ell^2 \bar{\square}) [\bar{g}_{ab} \bar{\square} - \bar{\nabla}_a \bar{\nabla}_b] R^{(1)} \right] = 0$$

Bianchi identity $\nabla^a \mathcal{E}_{ab} = 0$ leads to

$$\begin{aligned} f_2(\ell^2 \bar{\square}) &= \frac{D-2}{2D} \left[f_1(\ell^2 \bar{\square}) - f_1(\ell^2 \bar{\square} + 2D\chi) \right] \\ &+ \frac{D-1}{D\ell^2} \left[f_3(\ell^2 \bar{\square} + 2D\chi) (\ell^2 \bar{\square} - \chi D) - f_3(\ell^2 \bar{\square}) \ell^2 \bar{\square} \right] \end{aligned}$$

Connection with the Lagrangian?

Connection with the Lagrangian?

These equations can be derived from a general quadratic Lagrangian

$$\mathcal{L}_{\text{eff}} = \lambda \left[\frac{(D-1)(D-2)}{\ell^2} + R + \ell^2 R F_1(\ell^2 \square) R + \ell^2 R_{ab} F_2(\ell^2 \square) R^{ab} + \ell^2 R_{abcd} F_3(\ell^2 \square) R^{abcd} \right]$$

$f_{1,2,3}$ = functions of $F_{1,2,3}$. Map is complicated for $\Lambda \neq 0$

Connection with the Lagrangian?

These equations can be derived from a general quadratic Lagrangian

$$\mathcal{L}_{\text{eff}} = \lambda \left[\frac{(D-1)(D-2)}{\ell^2} + R + \ell^2 R F_1(\ell^2 \bar{\square}) R + \ell^2 R_{ab} F_2(\ell^2 \bar{\square}) R^{ab} + \ell^2 R_{abcd} F_3(\ell^2 \bar{\square}) R^{abcd} \right]$$

$f_{1,2,3}$ = functions of $F_{1,2,3}$. Map is complicated for $\Lambda \neq 0$

For a flat background we get

$$f_1 = \frac{\lambda}{2} \left[1 + \left[4F_3(\ell^2 \bar{\square}) + F_2(\ell^2 \bar{\square}) \right] \ell^2 \bar{\square} \right]$$

$$f_2 = 0$$

$$f_3 = -\frac{\lambda}{2} \ell^2 \left[2F_1(\ell^2 \bar{\square}) + F_2(\ell^2 \bar{\square}) + 2F_3(\ell^2 \bar{\square}) \right]$$

Degrees of freedom

Trace of the equations:

$$-\frac{\lambda}{4} \left[(D-2) - \ell^2 \bar{\square} \left[4F_3(\ell^2 \bar{\square}) + DF_2(\ell^2 \bar{\square}) + 4(D-1)F_1(\ell^2 \bar{\square}) \right] \right] R^{(1)} = 0$$

There are **scalar modes** unless $4F_3 + DF_2 + 4(D-1)F_1 = 0$

Degrees of freedom

Trace of the equations:

$$-\frac{\lambda}{4} \left[(D-2) - \ell^2 \bar{\square} \left[4F_3(\ell^2 \bar{\square}) + DF_2(\ell^2 \bar{\square}) + 4(D-1)F_1(\ell^2 \bar{\square}) \right] \right] R^{(1)} = 0$$

There are **scalar modes** unless $4F_3 + DF_2 + 4(D-1)F_1 = 0$

Traceless part is given by

$$\frac{\lambda}{2} \left\{ \left[1 + \left[4F_3(\ell^2 \bar{\square}) + F_2(\ell^2 \bar{\square}) \right] \ell^2 \bar{\square} \right] R^{(1)}_{\langle ab \rangle} - \ell^2 \left[2F_1(\ell^2 \bar{\square}) + F_2(\ell^2 \bar{\square}) + 2F_3(\ell^2 \bar{\square}) \right] \bar{\nabla}_{\langle a} \bar{\nabla}_{b \rangle} R^{(1)} \right\} = 0$$

Massive gravitons: roots of $1 + \left[4F_3(\ell^2 \bar{\square}) + F_2(\ell^2 \bar{\square}) \right] \ell^2 \bar{\square} \equiv \mathcal{F}(\bar{\square})$

Degrees of freedom

Trace of the equations:

$$-\frac{\lambda}{4} \left[(D-2) - \ell^2 \bar{\square} \left[4F_3(\ell^2 \bar{\square}) + DF_2(\ell^2 \bar{\square}) + 4(D-1)F_1(\ell^2 \bar{\square}) \right] \right] R^{(1)} = 0$$

There are **scalar modes** unless $4F_3 + DF_2 + 4(D-1)F_1 = 0$

Traceless part is given by

$$\frac{\lambda}{2} \left\{ \left[1 + \left[4F_3(\ell^2 \bar{\square}) + F_2(\ell^2 \bar{\square}) \right] \ell^2 \bar{\square} \right] R_{\langle ab \rangle}^{(1)} - \ell^2 \left[2F_1(\ell^2 \bar{\square}) + F_2(\ell^2 \bar{\square}) + 2F_3(\ell^2 \bar{\square}) \right] \bar{\nabla}_{\langle a} \bar{\nabla}_{b \rangle} R^{(1)} \right\} = 0$$

Massive gravitons: roots of $1 + \left[4F_3(\ell^2 \bar{\square}) + F_2(\ell^2 \bar{\square}) \right] \ell^2 \bar{\square} \equiv \mathcal{F}(\bar{\square})$

Possible resolutions:

- Nonlocal gravity: $\mathcal{F}(\square)$ is an entire functions with no zeroes [Mazumdar+...]
- Einstein-like gravities: $4F_3(\ell^2 \bar{\square}) + F_2(\ell^2 \bar{\square}) = 0$ [Bueno, PAC]

- 1 HIGHER-DERIVATIVE GRAVITY WITH COVARIANT DERIVATIVES
- 2 **INDUCED GRAVITY ON THE BRANE**
- 3 NONLOCAL MASSIVE GRAVITY IN $D = 3$
- 4 CONCLUSIONS

Consider Einstein gravity in AdS_{D+1}

$$I_{\text{Einstein}} = \frac{1}{16\pi G_{D+1}} \left[\int_{\mathcal{M}} d^{D+1}x \sqrt{-\hat{g}} \left(\hat{R} + \frac{D(D-1)}{\ell^2} \right) + 2 \int_{\partial\mathcal{M}} d^Dx \sqrt{-g} K \right]$$

INDUCED GRAVITY ON THE BRANE

Consider Einstein gravity in AdS_{D+1}

$$I_{\text{Einstein}} = \frac{1}{16\pi G_{D+1}} \left[\int_{\mathcal{M}} d^{D+1}x \sqrt{-\hat{g}} \left(\hat{R} + \frac{D(D-1)}{\ell^2} \right) + 2 \int_{\partial\mathcal{M}} d^Dx \sqrt{-g} K \right]$$

We add a brane at a cutoff $\rho = \ell/\epsilon$ before the AdS boundary. We have

$$I_{\text{Einstein}} + I_{\text{brane}} = I_{\text{bgrav}} + I_{\text{CFT}}$$

$$I_{\text{brane}} = -T \int_{\rho=\ell/\epsilon} d^Dx \sqrt{-g}, \quad I_{\text{bgrav}} = \frac{1}{16\pi G_{D+1}} \int d^Dx \sqrt{-g} \mathcal{L}$$

INDUCED GRAVITY ON THE BRANE

Consider Einstein gravity in AdS_{D+1}

$$I_{\text{Einstein}} = \frac{1}{16\pi G_{D+1}} \left[\int_{\mathcal{M}} d^{D+1}x \sqrt{-\hat{g}} \left(\hat{R} + \frac{D(D-1)}{\ell^2} \right) + 2 \int_{\partial\mathcal{M}} d^Dx \sqrt{-g} K \right]$$

We add a brane at a cutoff $\rho = \ell/\epsilon$ before the AdS boundary. We have

$$I_{\text{Einstein}} + I_{\text{brane}} = I_{\text{bgrav}} + I_{\text{CFT}}$$

$$I_{\text{brane}} = -T \int_{\rho=\ell/\epsilon} d^Dx \sqrt{-g}, \quad I_{\text{bgrav}} = \frac{1}{16\pi G_{D+1}} \int d^Dx \sqrt{-g} \mathcal{L}$$

Variation with respect to the boundary metric yields

$$\Pi_{ab} = K_{ab} - g_{ab}(K - T)$$

where Π_{ab} = boundary stress-energy tensor, equivalent to the equations of motion of I_{bgrav}

$$\Pi^{ab} \equiv \frac{2}{\sqrt{-g}} \frac{\delta I_{\text{bgrav}}}{\delta g_{ab}}$$

Induced gravity [Kraus, Larsen, Siebelink]

- 1 Start with Einstein equation's in the bulk $\hat{G}_{\mu\nu} = \frac{D(D-1)}{\ell^2} \hat{g}_{\mu\nu}$

Induced gravity [Kraus, Larsen, Siebelink]

- 1 Start with Einstein equation's in the bulk $\hat{G}_{\mu\nu} = \frac{D(D-1)}{\ell^2} \hat{g}_{\mu\nu}$
- 2 Evaluate them at the brane and decompose them in extrinsic and intrinsic curvature

Induced gravity [Kraus, Larsen, Siebelink]

- 1 Start with Einstein equation's in the bulk $\hat{G}_{\mu\nu} = \frac{D(D-1)}{\ell^2} \hat{g}_{\mu\nu}$
- 2 Evaluate them at the brane and decompose them in extrinsic and intrinsic curvature
- 3 Use the relation between K_{ab} and Π_{ab}

Induced gravity [Kraus, Larsen, Siebelink]

- 1 Start with Einstein equation's in the bulk $\hat{G}_{\mu\nu} = \frac{D(D-1)}{\ell^2} \hat{g}_{\mu\nu}$
- 2 Evaluate them at the brane and decompose them in extrinsic and intrinsic curvature
- 3 Use the relation between K_{ab} and Π_{ab}

Normal-normal component yields

$$\frac{1}{D-1}(\Pi + DT)^2 - (\Pi_{ab} + Tg_{ab})(\Pi^{ab} + Tg^{ab}) - R = \frac{D(D-1)}{\ell^2}$$

The tension fixes the cosmological constant on the brane

Induced gravity [Kraus, Larsen, Siebelink]

- 1 Start with Einstein equation's in the bulk $\hat{G}_{\mu\nu} = \frac{D(D-1)}{\ell^2} \hat{g}_{\mu\nu}$
- 2 Evaluate them at the brane and decompose them in extrinsic and intrinsic curvature
- 3 Use the relation between K_{ab} and Π_{ab}

Normal-normal component yields

$$\frac{1}{D-1}(\Pi + DT)^2 - (\Pi_{ab} + Tg_{ab})(\Pi^{ab} + Tg^{ab}) - R = \frac{D(D-1)}{\ell^2}$$

The tension fixes the cosmological constant on the brane

Choosing $T = (D-1)/\ell$ we get $\Lambda = 0$

$$\Pi = \frac{\ell}{2} \left[R + \Pi_{ab}\Pi^{ab} - \frac{1}{D-1}\Pi^2 \right]$$

This identity **fixes the induced gravity theory**

In order to derive the form of the equations, we expand them in derivatives

$$\mathcal{L} = \sum_{n=1}^{\infty} \ell^{2n-1} \mathcal{L}_{(n)}, \quad \Pi^{ab} = \sum_{n=1}^{\infty} \ell^{2n-1} \Pi_{(n)}^{ab}$$

In order to derive the form of the equations, we expand them in derivatives

$$\mathcal{L} = \sum_{n=1}^{\infty} \ell^{2n-1} \mathcal{L}_{(n)}, \quad \Pi^{ab} = \sum_{n=1}^{\infty} \ell^{2n-1} \Pi_{(n)}^{ab}$$

Then, we get

$$\begin{aligned} \Pi_{(1)} &= \frac{R}{2} \\ \Pi_{(n)} &= \frac{1}{2} \sum_{i=1}^{n-1} \left[\Pi_{(i)ab} \Pi_{(n-i)}^{ab} - \frac{1}{D-1} \Pi_{(i)} \Pi_{(n-i)} \right], \quad n \geq 2 \end{aligned}$$

In order to derive the form of the equations, we expand them in derivatives

$$\mathcal{L} = \sum_{n=1}^{\infty} \ell^{2n-1} \mathcal{L}_{(n)}, \quad \Pi^{ab} = \sum_{n=1}^{\infty} \ell^{2n-1} \Pi_{(n)}^{ab}$$

Then, we get

$$\begin{aligned} \Pi_{(1)} &= \frac{R}{2} \\ \Pi_{(n)} &= \frac{1}{2} \sum_{i=1}^{n-1} \left[\Pi_{(i)ab} \Pi_{(n-i)}^{ab} - \frac{1}{D-1} \Pi_{(i)} \Pi_{(n-i)} \right], \quad n \geq 2 \end{aligned}$$

The other ingredient we need to solve this recursive relation is

$$\Pi_{(n)} = (D - 2n) \mathcal{L}_{(n)} + \text{total derivative}$$

→ We get $\mathcal{L}_{(n)}$ from the trace of the equation of motion $\Pi_{(n)}$

In order to derive the form of the equations, we expand them in derivatives

$$\mathcal{L} = \sum_{n=1}^{\infty} \ell^{2n-1} \mathcal{L}_{(n)}, \quad \Pi^{ab} = \sum_{n=1}^{\infty} \ell^{2n-1} \Pi_{(n)}^{ab}$$

Then, we get

$$\begin{aligned} \Pi_{(1)} &= \frac{R}{2} \\ \Pi_{(n)} &= \frac{1}{2} \sum_{i=1}^{n-1} \left[\Pi_{(i)ab} \Pi_{(n-i)}^{ab} - \frac{1}{D-1} \Pi_{(i)} \Pi_{(n-i)} \right], \quad n \geq 2 \end{aligned}$$

The other ingredient we need to solve this recursive relation is

$$\Pi_{(n)} = (D - 2n) \mathcal{L}_{(n)} + \text{total derivative}$$

→ We get $\mathcal{L}_{(n)}$ from the trace of the equation of motion $\Pi_{(n)}$

$$\mathcal{L}_{(1)} = \frac{R}{2(D-2)}, \quad \Pi_{(1)ab} = -\frac{1}{D-2} G_{ab}$$

Focus on the quadratic part of the Lagrangian

$$\mathcal{L}_{(n)} = \alpha_n R \square^{n-2} R + \beta_n R^{ab} \square^{n-2} R_{ab} + O(R^3)$$

Focus on the quadratic part of the Lagrangian

$$\mathcal{L}_{(n)} = \alpha_n R \square^{n-2} R + \beta_n R^{ab} \square^{n-2} R_{ab} + O(R^3)$$

Recurrence relation becomes

$$\alpha_n = -\frac{D}{4(D-1)}\beta_n, \quad \beta_n = \frac{2}{(D-2n)} \left[\frac{\beta_{n-1}}{(D-2)} + \sum_{i=2}^{n-2} \beta_i \beta_{n-i} \right]$$

with $\beta_2 = \frac{1}{2(D-2)^2(D-4)}$.

Focus on the quadratic part of the Lagrangian

$$\mathcal{L}_{(n)} = \alpha_n R \square^{n-2} R + \beta_n R^{ab} \square^{n-2} R_{ab} + O(R^3)$$

Recurrence relation becomes

$$\alpha_n = -\frac{D}{4(D-1)}\beta_n, \quad \beta_n = \frac{2}{(D-2n)} \left[\frac{\beta_{n-1}}{(D-2)} + \sum_{i=2}^{n-2} \beta_i \beta_{n-i} \right]$$

with $\beta_2 = \frac{1}{2(D-2)^2(D-4)}$. This can be reformulated in terms of a generating function

$$f(x) = \sum_{n=2}^{\infty} \beta_n x^{2n-D}$$

Focus on the quadratic part of the Lagrangian

$$\mathcal{L}_{(n)} = \alpha_n R \square^{n-2} R + \beta_n R^{ab} \square^{n-2} R_{ab} + O(R^3)$$

Recurrence relation becomes

$$\alpha_n = -\frac{D}{4(D-1)}\beta_n, \quad \beta_n = \frac{2}{(D-2n)} \left[\frac{\beta_{n-1}}{(D-2)} + \sum_{i=2}^{n-2} \beta_i \beta_{n-i} \right]$$

with $\beta_2 = \frac{1}{2(D-2)^2(D-4)}$. This can be reformulated in terms of a generating function

$$f(x) = \sum_{n=2}^{\infty} \beta_n x^{2n-D}$$

The recurrence relation becomes a differential equation

$$f'(x) = (4-D)\beta_2 x^{3-D} - \frac{2}{D-2} x f(x) - 2x^{D-1} f(x)^2$$

$$I_{\text{bgrav}}^{(2)} = \frac{1}{16\pi G_{D+1}} \int d^D x \sqrt{-g} \left[\frac{\ell R}{2(D-2)} + \sum_{n=2}^{\infty} \beta_n \ell^{2n-1} \left(R^{ab} \square^{n-2} R_{ab} - \frac{D}{4(D-1)} R \square^{n-2} R \right) \right]$$

$$I_{\text{bgrav}}^{(2)} = \frac{1}{16\pi G_D} \int d^D x \sqrt{-g} \left[R + \ell^2 R^{ab} F(\ell^2 \square) \left(R_{ab} - \frac{D}{4(D-1)} g_{ab} R \right) \right]$$

where

$$F(\ell^2 \square) = 2(D-2) \sum_{n=2}^{\infty} \beta_n (\ell^2 \square)^{n-2}$$

and $G_D = 2(D-2)G_{D+1}/\ell$.

$$I_{\text{bgrav}}^{(2)} = \frac{1}{16\pi G_D} \int d^D x \sqrt{-g} \left[R + \ell^2 R^{ab} F(\ell^2 \square) \left(R_{ab} - \frac{D}{4(D-1)} g_{ab} R \right) \right]$$

where

$$F(\ell^2 \square) = 2(D-2) \sum_{n=2}^{\infty} \beta_n (\ell^2 \square)^{n-2}$$

and $G_D = 2(D-2)G_{D+1}/\ell$.

$f(x) = \frac{1}{2(D-2)} x^{4-D} F(x^2) \Rightarrow F(x)$ satisfies the equation

$$F'(x) = (D-4) \frac{F(x) - F(0)}{2x} - \frac{1}{2(D-2)} (2F(x) + xF(x)^2)$$

where $F(0) = 2(D-2)\beta_2 = \frac{1}{(D-2)(D-4)}$

$$F_D(x) = \frac{D(D-2)}{x^2} - \frac{1}{x} - \frac{(D-2)Y_{\frac{D+2}{2}}(\sqrt{x})}{x^{3/2}Y_{\frac{D}{2}}(\sqrt{x})}$$

$$F_D(x) = \frac{D(D-2)}{x^2} - \frac{1}{x} - \frac{(D-2)Y_{\frac{D+2}{2}}(\sqrt{x})}{x^{3/2}Y_{\frac{D}{2}}(\sqrt{x})}$$

The theory is well-defined in odd dimensions

$$F_3(x) = -\frac{\sin(\sqrt{x})}{x \sin(\sqrt{x}) + \sqrt{x} \cos(\sqrt{x})} \approx -1 + \frac{2x}{3} - \frac{7x^2}{15} + \frac{34x^3}{105} + \dots$$

$$F_5(x) = \frac{\cos(\sqrt{x})}{(3-x) \cos(\sqrt{x}) + 3\sqrt{x} \sin(\sqrt{x})} \approx \frac{1}{3} - \frac{2x}{9} + \frac{x^2}{27} + \frac{2x^3}{405} + \dots$$

$$F_7(x) = -\frac{\sqrt{x} \sin(\sqrt{x}) + \cos(\sqrt{x})}{(x-15)\sqrt{x} \sin(\sqrt{x}) + 3(2x-5) \cos(\sqrt{x})} \approx \frac{1}{15} + \frac{2x}{75} - \frac{13x^2}{1125} + \frac{22x^3}{16875} + \dots$$

$$F_D(x) = \frac{D(D-2)}{x^2} - \frac{1}{x} - \frac{(D-2)Y_{\frac{D+2}{2}}(\sqrt{x})}{x^{3/2}Y_{\frac{D}{2}}(\sqrt{x})}$$

The theory is well-defined in odd dimensions

$$F_3(x) = -\frac{\sin(\sqrt{x})}{x \sin(\sqrt{x}) + \sqrt{x} \cos(\sqrt{x})} \approx -1 + \frac{2x}{3} - \frac{7x^2}{15} + \frac{34x^3}{105} + \dots$$

$$F_5(x) = \frac{\cos(\sqrt{x})}{(3-x)\cos(\sqrt{x}) + 3\sqrt{x}\sin(\sqrt{x})} \approx \frac{1}{3} - \frac{2x}{9} + \frac{x^2}{27} + \frac{2x^3}{405} + \dots$$

$$F_7(x) = -\frac{\sqrt{x}\sin(\sqrt{x}) + \cos(\sqrt{x})}{(x-15)\sqrt{x}\sin(\sqrt{x}) + 3(2x-5)\cos(\sqrt{x})} \approx \frac{1}{15} + \frac{2x}{75} - \frac{13x^2}{1125} + \frac{22x^3}{16875} + \dots$$

Singularities in even dimensions

$$F_4(x) = \frac{8}{x^2} - \frac{1}{x} - \frac{2Y_3(\sqrt{x})}{x^{3/2}Y_2(\sqrt{x})} \approx \frac{1}{4} [-2\gamma_E - \log(x/4)] + \frac{1}{8} [-1 + \gamma_E + \log(x/4)]x + \dots$$

Linearized equations

$$F_1 = F, F_2 = -D/(4(D-1))F, F_3 = 0$$

Trace of linearized equations

$$-\frac{(D-2)}{64\pi G}R^{(1)} = 0$$

No scalar modes!

Linearized equations

$$F_1 = F, F_2 = -D/(4(D-1))F, F_3 = 0$$

Trace of linearized equations

$$-\frac{(D-2)}{64\pi G} R^{(1)} = 0$$

No scalar modes!

The linearized equations reduce to

$$\frac{1}{32\pi G} [1 + F(\ell^2 \bar{\square}) \ell^2 \bar{\square}] G_{ab}^{(1)} = 0$$

Linearized equations

$$F_1 = F, F_2 = -D/(4(D-1))F, F_3 = 0$$

Trace of linearized equations

$$-\frac{(D-2)}{64\pi G}R^{(1)} = 0$$

No scalar modes!

The linearized equations reduce to

$$-\frac{1}{64\pi G} \left[1 + F \left(\ell^2 \bar{\square} \right) \ell^2 \bar{\square} \right] \bar{\square} h_{ab} = 0, \quad h = 0, \quad \nabla^a h_{ab} = 0$$

Linearized equations

$$F_1 = F, F_2 = -D/(4(D-1))F, F_3 = 0$$

Trace of linearized equations

$$-\frac{(D-2)}{64\pi G}R^{(1)} = 0$$

No scalar modes!

The linearized equations reduce to

$$-\frac{1}{64\pi G} \left[1 + F \left(\ell^2 \bar{\square} \right) \ell^2 \bar{\square} \right] \bar{\square} h_{ab} = 0, \quad h = 0, \quad \nabla^a h_{ab} = 0$$

Propagator

$$P_D(k) = \frac{64\pi G_D}{(D-2)} \left[\frac{i\ell k Y_{\frac{D+2}{2}}(i\ell k)}{Y_{\frac{D}{2}}(i\ell k)} - D \right]^{-1}$$

Poles of the propagator represent propagating modes with mass $m^2 = -k^2$

Five dimensions

$$\frac{P_5(k)}{64\pi G_5} = \frac{1}{\ell^2 k^2} + \frac{1}{3 - 3\ell k \tanh(\ell k)}$$

INDUCED GRAVITY ON THE BRANE

Five dimensions

$$\frac{P_5(k)}{64\pi G_5} = \frac{1}{\ell^2 k^2} + \frac{1}{3 - 3\ell k \tanh(\ell k)}$$

Infinitely massive modes

$$\Delta m_n \sim \frac{\pi}{\ell}$$

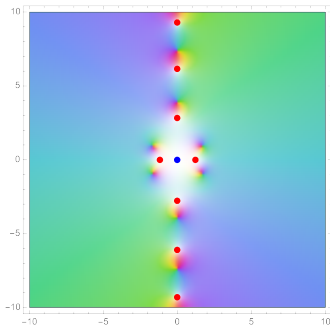
And one tachyonic mode

$$m_t^2 \approx -\frac{1.43923}{\ell^2}$$

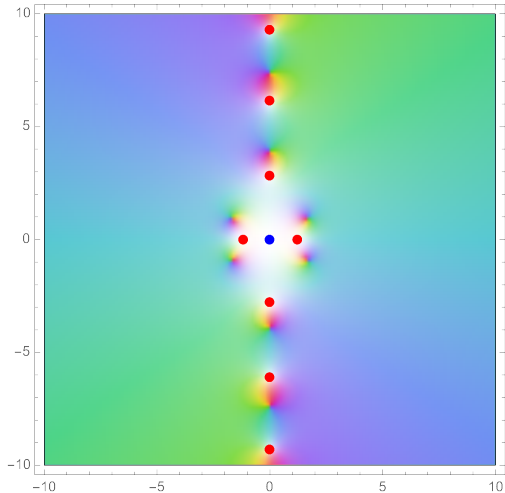
All modes have negative kinetic energy!

$$\frac{P_5(k^2 \rightarrow -m_j^2)}{64\pi G_5} = -\frac{2}{3\ell^2[k^2 + m_j^2]}$$

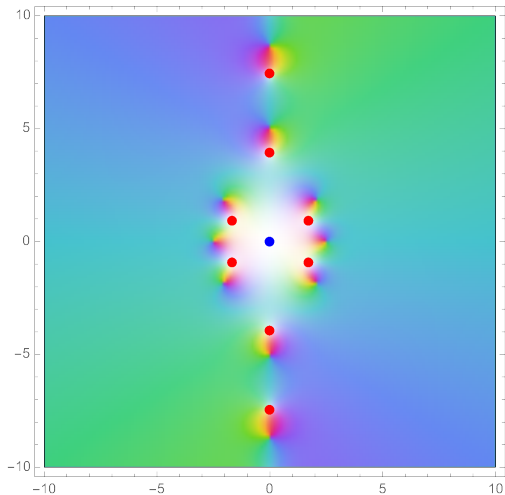
Can the coupling to the CFT alleviate this “pathological” behavior?



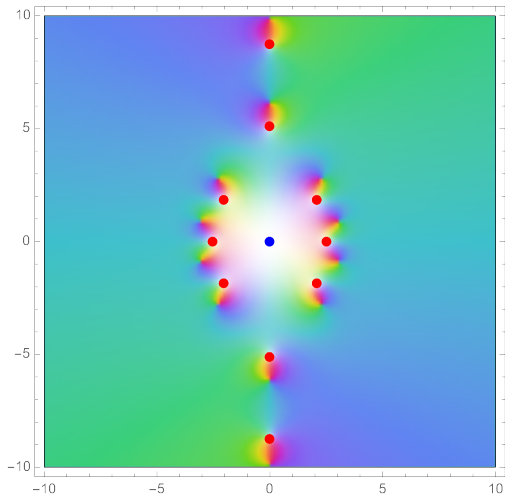
Higher dimensions: $D = 5$



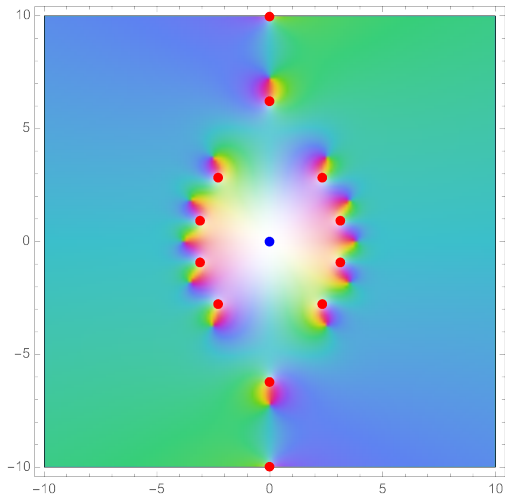
Higher dimensions: $D = 7$



Higher dimensions: $D = 9$



Higher dimensions: $D = 11$



Three dimensional case is special

Three dimensional case is special

$$\frac{P_3(k)}{64\pi G_3} = \frac{1}{\ell^2 k^2} - \frac{\ell k \tanh(\ell k)}{\ell^2 k^2}$$

Three dimensional case is special

$$\frac{P_3(k)}{64\pi G_3} = \frac{1}{\ell^2 k^2} - \frac{\ell k \tanh(\ell k)}{\ell^2 k^2}$$

$$m_n = \frac{\pi}{\ell} \left(n + \frac{1}{2} \right), \quad n = 0, 1, 2, \dots$$

No tachyons! infinite tower of equally-spaced massive modes

Three dimensional case is special

$$\frac{P_3(k)}{64\pi G_3} = \frac{1}{\ell^2 k^2} - \frac{\ell k \tanh(\ell k)}{\ell^2 k^2}$$

$$m_n = \frac{\pi}{\ell} \left(n + \frac{1}{2} \right), \quad n = 0, 1, 2, \dots$$

No tachyons! infinite tower of equally-spaced massive modes

These are still ghosts...

...however, the Einsteinian mode $k^2 = 0$ in $D = 3$ is **non-propagating**

Three dimensional case is special

$$\frac{P_3(k)}{64\pi G_3} = \frac{1}{\ell^2 k^2} - \frac{\ell k \tanh(\ell k)}{\ell^2 k^2}$$

$$m_n = \frac{\pi}{\ell} \left(n + \frac{1}{2} \right), \quad n = 0, 1, 2, \dots$$

No tachyons! infinite tower of equally-spaced massive modes

These are still ghosts...

...however, the Einsteinian mode $k^2 = 0$ in $D = 3$ is **non-propagating**

The theory has negative-definite energy!

$$\mathcal{L} \rightarrow -\mathcal{L}$$

- 1 HIGHER-DERIVATIVE GRAVITY WITH COVARIANT DERIVATIVES
- 2 INDUCED GRAVITY ON THE BRANE
- 3 NONLOCAL MASSIVE GRAVITY IN $D = 3$**
- 4 CONCLUSIONS

We define NLMG as the 3D braneworld gravity with a global sign change

$$I_{\text{NLMG}} = \frac{1}{16\pi G} \int d^3x \sqrt{-g} \left[-R + \ell^2 R^{ab} F(\ell^2 \square) \left(R_{ab} - \frac{3}{8} g_{ab} R \right) \right]$$

$$F(x) \equiv \frac{\sin(\sqrt{x})}{x \sin(\sqrt{x}) + \sqrt{x} \cos(\sqrt{x})}$$

We define NLMG as the 3D braneworld gravity with a global sign change

$$I_{\text{NLMG}} = \frac{1}{16\pi G} \int d^3x \sqrt{-g} \left[-R + \ell^2 R^{ab} F(\ell^2 \square) \left(R_{ab} - \frac{3}{8} g_{ab} R \right) \right]$$

$$F(x) \equiv \frac{\sin(\sqrt{x})}{x \sin(\sqrt{x}) + \sqrt{x} \cos(\sqrt{x})}$$

- It generalizes New Massive Gravity [[Bergshoeff](#), [Hohm](#), [Townsend](#)]

$$\mathcal{L} = -R + \ell^2 \left(R^{ab} R_{ab} - \frac{3}{8} R^2 \right) + \mathcal{O}(\ell^4)$$

- It propagates an infinite tower of positive-energy massive gravitons
- Dynamical gravity in $D = 3$

Newtonian potential

$$ds^2 = -(1 + 2\Phi) dt^2 + (1 - 2\Psi) d\vec{x}^2$$

Assuming a pressureless stress energy tensor with energy density $\rho = T_{tt}$, these potentials satisfy the equations

$$\begin{aligned}\square\Psi + \frac{1}{2}F\left(\ell^2\square\right)\ell^2\square^2(\Phi + \Psi) &= -8\pi G\rho \\ \square(\Psi - \Phi) &= -8\pi G\rho\end{aligned}$$

Newtonian potential

$$ds^2 = -(1 + 2\Phi) dt^2 + (1 - 2\Psi) d\vec{x}^2$$

Assuming a pressureless stress energy tensor with energy density $\rho = T_{tt}$, these potentials satisfy the equations

$$\begin{aligned}\square\Psi + \frac{1}{2}F\left(\ell^2\square\right)\ell^2\square^2(\Phi + \Psi) &= -8\pi G\rho \\ \square(\Psi - \Phi) &= -8\pi G\rho\end{aligned}$$

Point-like mass:

$$\Psi - \Phi \propto \log(r/M) \quad \Rightarrow \text{large gauge transformation}$$

For Einstein gravity $\Phi_{\text{EG}} = 0$ and the metric is locally flat

Newtonian potential

$$ds^2 = -(1 + 2\Phi) dt^2 + (1 - 2\Psi) d\vec{x}^2$$

Assuming a pressureless stress energy tensor with energy density $\rho = T_{tt}$, these potentials satisfy the equations

$$\begin{aligned}\square\Psi + \frac{1}{2}F\left(\ell^2\square\right)\ell^2\square^2(\Phi + \Psi) &= -8\pi G\rho \\ \square(\Psi - \Phi) &= -8\pi G\rho\end{aligned}$$

Point-like mass:

$$\Psi - \Phi \propto \log(r/M) \quad \Rightarrow \text{large gauge transformation}$$

For Einstein gravity $\Phi_{\text{EG}} = 0$ and the metric is locally flat

For the massive gravities:

$$\begin{aligned}\Phi_{\text{NMG}} &= -4\pi G M K_0(mr) \\ \Phi_{\text{NLMG}} &= -4\pi G \ell M \int_0^\infty dk J_0(kr) \tanh(\ell k)\end{aligned}$$

Asymptotic behavior

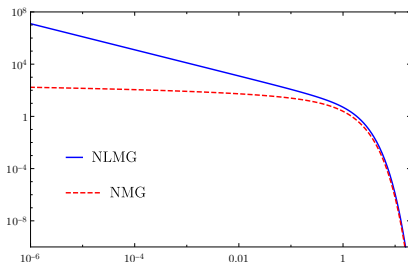
$$\Phi_{\text{NMG}} \sim r^{-1/2} e^{-mr}$$

$$\Phi_{\text{NLMG}} \sim r^{-1/2} e^{-m_1 r}$$

Small r behavior

$$\Phi_{\text{NMG}} \sim \log(r)$$

$$\Phi_{\text{NLMG}} \sim -\frac{4\pi G\ell M}{r}$$



- **We recover the $1/r$ behavior of 4 dimensions!**
- **Genuine nonlocal effect:** cannot be obtained with a finite number of derivatives
- Could there be black hole solutions?

- 1 HIGHER-DERIVATIVE GRAVITY WITH COVARIANT DERIVATIVES
- 2 INDUCED GRAVITY ON THE BRANE
- 3 NONLOCAL MASSIVE GRAVITY IN $D = 3$
- 4 CONCLUSIONS

- Induced gravity on the brane is an infinite-derivative (nonlocal) gravity
- These theories are not “healthy”: negative energy modes. What is the meaning of this?
- Do the massive modes agree with the Kaluza-Klein modes (in the Randall-Sundrum sense)?
- In 3D, the braneworld gravity with reversed sign is a well-behaved theory which generalizes NMG. What other aspects of the theory could we study?
- Extensions: spectrum of braneworld theories in AdS/dS, Lovelock braneworlds.

- Induced gravity on the brane is an infinite-derivative (nonlocal) gravity
- These theories are not “healthy”: negative energy modes. What is the meaning of this?
- Do the massive modes agree with the Kaluza-Klein modes (in the Randall-Sundrum sense)?
- In 3D, the braneworld gravity with reversed sign is a well-behaved theory which generalizes NMG. What other aspects of the theory could we study?
- Extensions: spectrum of braneworld theories in AdS/dS, Lovelock braneworlds.

Thank you for your attention!